ORCA – Online Research @ Cardiff



This is an Open Access document downloaded from ORCA, Cardiff University's institutional repository:https://orca.cardiff.ac.uk/id/eprint/99341/

This is the author's version of a work that was submitted to / accepted for publication.

Citation for final published version:

Nanne, Josefine A. M., Millet, Marc-Alban, Burton, Kevin W., Dale, Chris W., Nowell, Geoff M. and Williams, Helen M. 2017. High precision osmium stable isotope measurements by double spike MC-ICP-MS and N-TIMS. Journal of Analytical Atomic Spectrometry 32 (4), pp. 749-765. 10.1039/C6JA00406G

Publishers page: http://dx.doi.org/10.1039/C6JA00406G

Please note:

Changes made as a result of publishing processes such as copy-editing, formatting and page numbers may not be reflected in this version. For the definitive version of this publication, please refer to the published source. You are advised to consult the publisher's version if you wish to cite this paper.

This version is being made available in accordance with publisher policies. See http://orca.cf.ac.uk/policies.html for usage policies. Copyright and moral rights for publications made available in ORCA are retained by the copyright holders.



1 2 3 4 5	High precision osmium stable isotope measurements by double spike MC-ICP-MS and N-TIMS.	
	Josefine A.M. Nanne ^{*a} , Marc-Alban Millet ^{a,b} , Kevin W. Burton ^a , Chris W. Dale ^a , Geoff M. Nowell ^a and Helen M. Williams ^{a,c}	
6	* Corresponding author. E-mail: j.a.nanne@durham.ac.uk	
7	a.	Department of Earth Sciences,
8		Durham University,
9		Science Labs, South Road,
10		Durham DH1 3LE, United Kingdom.
11		
12	b.	Present address: School of Earth and Ocean Sciences,
13		Cardiff University,
14		Main building, Park Place,
15		Cardiff CF10 3AT, United Kingdom.
16		
17	с.	Present address: Department of Earth Sciences,
18		The University of Cambridge,
19		Downing Street,
20		Cambridge CB2 3EQ, United Kingdom.
21		

22 ABSTRACT

23 Osmium stable isotopes provide a new, potentially powerful tool with which to investigate a diverse 24 range of geological processes including planetary formation, ore-genesis and weathering. In this paper, 25 we present a new technique for high precision measurement of osmium (Os) stable isotope ratios by 26 both Multiple-Collector Inductively Coupled Plasma Mass Spectrometry (MC-ICP-MS) and Negative ion 27 Thermal Ionisation Mass Spectrometry (N-TIMS). We use a 1880s-1900s double spike, composed of 28 61% 188Os and 39% 190Os, to correct for mass dependent fractionation resulting from sample 29 preparation and isotope measurement, with the ideal spike to sample ratio being 55:45. Isotope ratios 30 are expressed as δ 1900s which is the per mil deviation in the measured 1900s/1880s ratio relative to 31 isotope reference material DROsS. Repeated analyses of double spiked DROsS for both MC-ICP-MS (n= 32 80 cycles) and N-TIMS (n= 280 cycles) show that an internal precision of 0.01-0.02 % on δ 1900s (2 se) 33 can be attained, with a long-term reproducibility of 0.016 ‰ and 0.029 ‰ (2 sd; n = 91 and 83, 34 respectively). The better reproducibility on MC-ICP-MS than on N-TIMS is, predominantly, due to 35 measurement at higher beam intensities (11-18 V with consumption of ~200 ng natural Os vs. 2-18 V 36 with consumption of 2.3 – 45 ng natural Os, respectively). In addition to stable isotope compositions, 37 our method allows for simultaneous measurement of 1870s/1880s and 1860s/1880s ratios with a 38 precision of <40 ppm (2 se; 80 cycles for MC-ICP-MS and 280 cycles for N-TIMS) and an external 39 reproducibility of 123-268 ppm and 234-361 ppm (2 sd; n= 91 for MC-ICP-MS and n= 83 for N-TIMS), 40 respectively. We demonstrate that a similar precision and reproducibility can be obtained for other 41 pure Os solutions as well as for geological materials. In addition, a range of analytical tests evaluates 42 and demonstrates the robustness of our method with regards to residual matrix effects and 43 interference correction, signal intensity and on-peak zero on MC-ICP-MS, and the effect of oxygen 44 corrections and isobaric interference on N-TIMS. Finally, we report the first Os stable isotope 45 compositions for geological reference materials, including mantle peridotites and chromitites, and one 46 ordinary chondrite.

48 **1. INTRODUCTION**

49 Osmium (Os) has two radiogenic isotopes (¹⁸⁶Os and ¹⁸⁷Os) and five naturally occurring stable isotopes: ¹⁸⁴Os, ¹⁸⁸Os, ¹⁸⁹Os, ¹⁹⁰Os, and ¹⁹²Os, with relative abundances of 0.02 %, 13.21 %, 16.11 %, 26.21 %, and 50 51 40.74 %, respectively. Osmium is a refractory and highly siderophile (Fe-loving) element and is therefore 52 of considerable interest in the study of planetary differentiation and formation of Earth's metallic core. 53 It is also a chalcophile (S-loving) element, and strongly partitions into sulphides. Consequently, Os 54 behaves compatibly during mantle melting, where sulphide remains as a residual phase in the source. 55 Taking advantage of the chemical properties of Os and the differences in behaviour between Re, Pt and 56 Os, the ¹⁸⁷Re-¹⁸⁷Os and ¹⁹⁰Pt-¹⁸⁶Os decay systems have become important chronometers and tracers in 57 both high and low-temperature geochemistry eg.1.2. More specifically, these radiogenic isotope systems 58 have been used to study early solar system dynamics e.g. 3, 4, planetary differentiation processes e.g. 5, 6, 59 mantle heterogeneity e.g. 7, 8, crustal growth and recycling e.g. 9-11, economic mineralization in ore deposits 60 e.g. 12, and the nature of weathering processes associated with brief climatic excursions e.g. 13-15. However, 61 despite the potential to use stable Os isotopes to investigate these same processes, thus far this system 62 remains unexplored.

A potential limitation for the use of Os stable isotopes in geochemistry is the small range of natural variation expected for high-mass elements. This is compounded by the very low abundance of Os in most terrestrial samples. However, recent advances in mass spectrometry and development of new techniques for stable isotope measurement have led to the discovery of significant and systematic stable isotope fractionation for high-mass stable isotope systems in both high and low-temperature environments¹⁶⁻²². This suggests that Os stable isotope variations may be measurable if a suitable high precision analytical method can be developed.

Additional complications that must be addressed in order to successfully measure Os stable isotopes include the non-quantitative recovery of Os during sample processing (typically on 60-80 %^{23,} and instrumental mass bias (MC-ICP-MS) or mass fractionation (N-TIMS) during measurement. Such fractionation can be overcome by the use of a double spike (DS) ²⁵. This approach has been shown to reliably account for mass-dependent stable isotope fractionation that can occur during all steps of sample processing (digestion, chemical separation and mass spectrometry) ^{18, 26, 27}.

76 In this paper, we present a new method for the measurement of high-precision stable Os 77 isotope ratios using a ¹⁸⁸Os-¹⁹⁰Os double spike by both plasma source (MC-ICP-MS) and negative 78 thermal ionisation mass spectrometry (N-TIMS). Method development on both machines allows for the 79 analysis of sample materials over a broad range of Os concentrations (>1 ppb) at precisions of 0.01-0.02 80 ‰ on δ^{190} Os (2 se; 80 cycles for MC-ICP-MS and 280 cycles for N-TIMS). In addition, we show that the method allows for simultaneous collection of stable isotopes and radiogenic isotope ratios, ¹⁸⁷Os/¹⁸⁸Os 81 82 and ¹⁸⁶Os/¹⁸⁸Os, at a precision of <40 ppm (2 se; 80 cycles for MC-ICP-MS and 280 cycles for N-TIMS). 83 We show the robustness, precision and accuracy of our method through a range of analytical tests and

repeated measurements of pure Os solutions and geological materials. Finally, data is presented for arange of geological materials.

86

87 2. OSMIUM DOUBLE SPIKE METHODOLOGY

88 2.1 Osmium Double Spike Design

89 The double spike (DS) approach requires four stable isotopes that are related to each other by massdependent stable isotope fractionation. For Os we can use ¹⁸⁸Os, ¹⁸⁹Os, ¹⁹⁰Os, and ¹⁹²Os. Departures 90 91 from mass-dependent fractionation, either due to cosmogenic effects or nucleosynthetic anomalies, 92 have not been detected in terrestrial samples but have been reported for extra-terrestrial materials. 93 Components of carbonaceous chondrites display mass independent Os isotopic anomalies although homogeneity is shown at the bulk meteoritic scale ^{e.g. 28, 29,30}. Mass independent anomalies at the bulk 94 95 rock scale have been shown for iron meteorites ^{e.g. 31} which should be considered when analysing such 96 meteorites.

97 The DS deconvolution used in this study is based on the geometric iterative resolution method 98 of Siebert et al. (2001)³². Measurements were also double checked using an algebraic resolution 99 method used by Millet and Dauphas (2014)³³ and Millet et al. (2016)³⁴ which yielded identical results. 100 Regardless of the approach, the DS deconvolution consists of resolving the following non-linear 101 equation:

102
$$R_m = \left[(1-f)R_{standard} (i_x/i_n)^{\alpha} + fR_{spike} \right] * (i_x/i_n)^{\beta}$$

103 where R_m , $R_{standard}$ and R_{spike} are the measured, standard and spike isotope ratios; *in* is the atomic weight 104 of the normalising isotope (¹⁸⁸Os); *ix* is the atomic weight of one of the three other isotopes used to 105 resolve the equation which in our method are ¹⁸⁹Os, ¹⁹⁰Os and ¹⁹²Os; *f* is the relative proportion of ¹⁸⁸Os 106 originating from the spike in the sample–spike mixture; α is the natural and β the processing and 107 instrumental exponential fractionation factors. In this study, reference material DROsS has been used 108 as standard. All Os stable isotope compositions are thus reported relative to DROsS, as the per mil 109 deviation (‰) of the ¹⁹⁰Os/¹⁸⁸Os ratios, here after reported as $\delta^{190}Os$:

110
$$\delta^{190/188} \text{Os}_{\text{DROsS}} = \left(\frac{\frac{190}{188} \text{Os}_{\text{sample}}}{\frac{190}{188} \text{Os}_{\text{DROsS}}} - 1\right) * 1000$$

Radiogenic isotopes ¹⁸⁶Os and ¹⁸⁷Os are not used in the DS deconvolution. Consequently, the spike proportion, and the geological and analytical fractionation factors resolved within the DS deconvolution can be used to calculate the ¹⁸⁶Os/¹⁸⁸Os and ¹⁸⁷Os/¹⁸⁸Os ratios. Osmium concentrations were determined by performing isotope dilution calculations.

115 The analytical uncertainty on double spike measurements is highly dependent on spike 116 composition and sample-spike mixing proportions. To establish the optimal composition and 117 proportions for Os, we have modelled the internal precision of a typical MC-ICP-MS measurement 118 following the model of Millet and Dauphas (2014)³³ which takes into account the errors associated to 119 Johnson noise and counting statistics (see ref. 33 for more details). This error model differs from that 120 of Rudge et al. (2009)³⁵ in that i) errors on the natural fractionation factor (alpha) are calculated on the 121 basis of a constant intensity for the most abundant isotope in the natural Os-DS mixture, rather than 122 calculated based on a constant total Os ion beam and; ii) determination of errors is done through Monte 123 Carlo modelling. In our model, the maximum intensity for the most abundant isotope was set at 6 V. A 124 measurement is assumed to comprise 80 cycles with an integration time of 8.3 seconds per cycle on 125 $10^{11}\Omega$ collectors at T=290 K. The model shows that a minimum internal error, of ~0.010 ‰ amu⁻¹ (2 se), 126 is acquired when using a 0.6:0.4 ¹⁸⁸Os – ¹⁹⁰Os mixed spike composition and a 0.55:0.45 spike-sample 127 mixture (Fig. 1). This is close to the optimal double spike composition as calculated by Rudge et al. 128 (2009)³⁵ of 0.66:0.34 ¹⁸⁸Os -¹⁹⁰Os for a 0.6:0.4 spike-sample mixture. It is important to note that the 2 129 se error is <0.020 ‰ amu⁻¹ when the sample proportion ranges between 15 % and 81 %. That errors 130 are minimal over a large range of spike-sample mixtures is of great utility when the Os concentration 131 of samples is not well known. A triple spike composition was considered in our calculations but was not 132 found to improve precision.

133

134 **2.2** Preparation and calibration of ¹⁸⁸Os – ¹⁹⁰Os double spike

Single ¹⁸⁸Os and ¹⁹⁰Os isotope spikes were purchased as fine-grained metallic powders from Trace Sciences International. In order to ensure purity of the double spike, the metal powders were individually digested using Carius tubes, followed by extraction and purification as described in section 3.2. Individual spikes were subsequently mixed in calculated optimal proportions and the resulting double spike was diluted with 3 M HCl to desired concentrations.

140 Calibration of the double spike isotope composition was achieved by measuring a pure 141 standard and pure double spike solution as well as a range of Durham Romil Osmium Standard (DROsS^{36,} 142 ³⁷)-DS mixtures. Practically, this involves iterative correction for instrumental mass fractionation of the pure DS measurements to generate a putative true DS composition, with an initial estimate provided 143 144 by the pure DROsS analysis. This putative true DS composition is then fed into the DS deconvolution 145 and used on all DS-DROsS mixtures, ranging from 0.1:0:9 to 09:0.1 mixing proportions. Calibration is 146 considered satisfactory once most mixtures, especially those around the optimum mixture proportions, 147 display a δ^{190} Os within analytical error of zero. For all these solutions, Carius tubes were used to ensure 148 standard-sample equilibration. Measurements were carried out by MC-ICP-MS with each analysis 149 comprising 220 cycles of 8.39 seconds and were all preceded by on-mass on-peak zero measurement 150 (20 integrations of 8.39 seconds). The isotope composition of the DS and of reference standard DROsS, 151 as used in the calibration and stable isotope calculations, can be found in Table 1.

153 3. MATERIALS AND SAMPLE PROCESSING

154 **3.1 Materials and samples**

Digestion and processing of standards and samples were carried out at Durham University. During the
 course of this study, we used Romil acids of variable grades (Super Purity Acids and Ultra Purity Acids)
 with blank levels suitable for the analytical procedure. The nitric acid was purified by sparging with
 filtered air. Acid dilutions were performed with ultrapure (18.2 MΩ·cm) water.

159 We use the Durham Romil Osmium Standard (DROsS)^{36, 37} as a reference standard. This is an international Os isotope reference material supplied by IAGeo Limited that has been widely distributed 160 161 e.g.38, 39 and for which Os isotope ratios have been determined to high precision and accuracy by both MC-ICP-MS³⁶ and N-TIMS³⁷. For method testing, we have also measured three in-house pure Os 162 163 reference solutions, ROMIL, SpecPure, and OsCaR, which were purchased from Romil Ltd, Alfa Aesar, 164 and provided by A. Poirier (OsCaR). As representative rock samples, we used well-characterized and commonly used reference materials: peridotites UB-N⁴⁰ and GP-13⁴¹, and chromitites CHR-Bkg^{42, 43} and 165 166 CHR-Pt+⁴⁴. Reference material GP-13 is an in-house PGE standard prepared at Durham University which 167 has been widely distributed. The other materials, UB-N, CHR-Bkg and CHR-Pt+, are distributed by CRPG-168 CRNS, Nancy, France. Furthermore, we included an ordinary H-chondrite, ZAG. These samples were 169 chosen to reflect the range of Os concentrations and sample matrices available for study.

170

171 **3.2** Sample digestion and chemical purification of Os

Samples were digested either by Carius tube (CT)^{23, 45} or high-pressure asher (HPA, Anton Paar)⁴⁴, with 172 173 respective maximum of 2.5 g and 2 g powder material per tube. Digestions were performed using 174 inverse aqua regia (1:2; 12 M HCl: 16 M HNO₃) for ≥16 h at temperatures of 230°C in the case of CT and 175 290-310°C for HPA digestions. The double spike was added to each tube prior to sample digestion. After 176 digestion, the extraction and purification of Os closely followed techniques described by Cohen et al. 177 (1996)²¹ and Birck et al. (1997)²². Separation of Os from the aqua regia digest solution was achieved by 178 solvent extraction using CHCl₃⁴⁶. Osmium was then back-extracted from the chloroform solution using 179 concentrated HBr and subsequently purified using the micro-distillation procedure of Roy-Barman $(1993)^{47}$ as described by Birck et al. $(1997)^{24}$. In the case of measurements by MC-ICP-MS, samples were 180 181 evaporated and chlorified three times using 300 μ L of 8 M HCl before being taken up in 3 M HCl to a 182 total Os concentration of ~0.5 µg mL⁻¹. When performing Os analyses by N-TIMS, samples were dried 183 down after microdistillation to approximately 1-2 μ L of sample solution and then loaded onto a Pt 184 ribbon filament. Typical total procedural blanks range between 0.01 and 0.10 pg Os which is similar to previous Os studies carried out at Durham University^{e.g. 48}. 185

186

187 4. MASS SPECTROMETRY

188 Osmium isotope ratios were measured at Durham University on a ThermoFisher Scientific Neptune MC-ICP-MS³⁶ and a Triton Plus N-TIMS^{37, 39, 49}. The use of both MC-ICP-MS and N-TIMS allowed us to exploit 189 190 the advantages of the different ionisation sources and provide a means to assess the relative accuracy 191 of each measurement. The overall introduction and ionisation efficiency of MC-ICP-MS is relatively low 192 ($\sim 0.08 \ \%^{36}$) which makes it better suited for analysing large samples and calibrating standards. The 193 advantage of the MC-ICP-MS is that Os is measured as Os⁺ which makes the mass spectrum relatively simple and corrections for isobaric interferences relatively straightforward. By contrast, N-TIMS has a 194 195 much higher ionisation efficiency for Os (1-5 %³⁷) which makes it ideal for samples with a low Os 196 abundance or when sample material is limited. However, Os is measured as the tri-oxide ion OsO3-197 which makes the mass spectrum complex and necessitates multiple oxide corrections that add to the 198 overall uncertainty. In both cases, internal errors are presented as the 2 se on δ^{190} Os of each integration, 199 once passed through a 2 se filter to remove outliers.

200

201 4.1 MC-ICP-MS

202 4.1.1 Instrument set-up and parameters MC-ICP-MS

203 The Neptune MC-ICP-MS was set-up for static simultaneous collection of all Os isotopes, apart from the least abundant ¹⁸⁴Os. We also collected ¹⁸³W, ¹⁸⁵Re, and ¹⁹⁴Pt in order to monitor and correct for isobaric 204 205 interference (Table 2). In this study, interference beam intensities were <0.05 mV for ¹⁸³W, <0.01 mV 206 for ¹⁸⁵Re, and <0.5 mV for ¹⁹⁴Pt. Typical instrument operating conditions were similar to those outlined 207 in Nowell et al. (2008)³⁶. Sample material was introduced in 3 M HCl using an Elemental Scientific 208 Incorporated (ESI) PFA-50 micro-flow nebuliser and Glass Expansion (GE) micro-cyclonic "Cinnabar" 209 spray-chamber. During typical run conditions, the measured sample uptake rate was ca. 80 µL min⁻¹ 210 with a sensitivity of ca. 30 V ppm⁻¹. Measurements were carried out in static mode in 1 block of 80 211 cycles of 8.389 seconds integration time each. Mass calibration was updated by peak-centering on the 212 centre-cup mass ¹⁸⁸Os at the start of each session and checked again at the end. Prior to every sample 213 run, baselines were obtained by on-mass on-peak zero (OPZ) measurements for 20 x 8.389 seconds integrations in clean 3 M HCl. Washouts between sample runs used 3 M HCl acid for at least 5 minutes 214 which resulted in the ¹⁸⁸Os beam decreasing below 2 mV (~0.03 % of the typical sample signal). Total 215 216 analysis time, including OPZ and wash-out, was approximately 15 minutes with consumption of 800 µL 217 sample solution (equivalent to 400 ng total Os in the case of a 0.5 μ g mL⁻¹ solution).

218

219 4.1.2 Data reduction MC-ICP-MS

Following analysis, all raw intensity data were exported and re-processed offline on a cycle-by-cycle basis using an in-house Excel spreadsheet. For MC-ICP-MS analyses, corrections were applied in the following order; baseline subtraction using the OPZ measurements, abundance sensitivity (1 ppm³⁶), isobaric interferences, and DS deconvolution. The abundance sensitivity is assumed to be similar to that
 determined by Nowell et al. (2008)³⁶.

225

226 4.1.2.1 Effect of beam intensity and on-peak zero

227 To evaluate the effect of beam size on the accuracy, the main 500 ppb DROsS-DS solution was diluted 228 to various lower concentrations (100 - 5 ppb total Os; Table 3). This resulted in total Os beam intensities 229 varying between 18 V and 0.1 V, with the corresponding ¹⁸⁸Os ranging between 6.7 V and 0.005 V. All analyses show Os stable isotope ratios as well as ¹⁸⁷Os/¹⁸⁸Os within analytical uncertainty of one 230 another (Fig. 2; Table 3). The ¹⁸⁶Os/¹⁸⁸Os only shows a significant deviation from the reference value 231 232 when the total Os beam intensity is 0.14 V, corresponding to solutions with concentrations of ~ 5 ppb 233 total Os which equals the consumption of ~4 ng total Os (i.e. natural plus DS) under standard running 234 conditions.

235 Analyses by MC-ICP-MS at low beam intensities critically rely on the absence of memory 236 effects. The on-peak zero (OPZ), used to account for baseline, can introduce an inaccuracy on the stable 237 isotope composition especially when measuring samples with very distinct isotope composition. The 238 importance of the wash-out and sample to sample memory has been discussed in detail by Nowell et 239 al. (2008)³⁶. They show that when certain precautions are taken (e.g., avoid drying out of the spray 240 chamber between sessions, keep standards and samples in the same chemical form), the effect on the 241 Os isotopic composition is negligible and should not prevent the acquisition of high precision data. 242 Considering that the natural variation in stable isotope compositions is expected to be small and the 243 proportion of sample to spike are well matched, the potential of memory effect on the stable isotope 244 ratios can, therefore, assumed to be minimal. In this study, the OPZ of the diluted DROsS solutions 245 display signal intensities between 0.04 % and 0.6 % relative to the peak Os signal (¹⁸⁸Os) in the main 246 run. The accuracy of the δ^{190} Os values even at low beam intensities suggests that the OPZ has an 247 insignificant effect on the accuracy at levels up to 0.6 %. For DROsS analyses run at total Os beam 248 intensities of >10 V, OPZ analyses are in general <2 mV and <0.04 % of the peak Os signal. The potential 249 of inaccurate δ^{190} Os values as a result of OPZ analyses can, therefore, be neglected. The fact that OPZ 250 correction is less accurate for ¹⁸⁶Os/¹⁸⁸Os is partly due to the lower intensities of ¹⁸⁶Os relative to all 251 other isotopes involved in the double-spike deconvolution (¹⁸⁸Os, ¹⁸⁹Os, ¹⁹⁰Os and ¹⁹²Os). The ¹⁸⁶Os 252 signal of the OPZ analyses at 0.14 V is 4 % of the main run beam. On-peak zero determination is, 253 therefore, critical for obtaining reliable 186 Os/ 188 Os simultaneously with δ^{190} Os by MC-ICP-MS.

254 It should be noted that the higher ionisation efficiency of the N-TIMS allows for small quantities 255 of Os to be measured at higher intensities. For example, while consumption of 4 ng total Os (5 ppb 256 solution) provides a total Os beam intensities of 0.1 V by MC-ICP-MS, we typically obtain a total Os 257 beam intensity of >1 V for a total Os load of \geq 5 ng by N-TIMS. Furthermore, Figure 7 shows that the 258 error is significantly increased when the ¹⁸⁸Os beam intensity drops below ~1 V (~2.7 V ^{total}Os) which is an important consideration to make when small variations in stable isotope composition are expected.

We, therefore, recommend that measurements on less than ~100 ng of total Os be performed by N-TIMS.

262

263 4.1.2.2 Matrix effects and interference corrections

264 Potential atomic isobaric inferences on Os isotopes can arise from the incomplete separation of Os from 265 W, Re and Pt. In addition, non-spectral matrix effects can create inaccurate results⁵⁰. To assess if our 266 method can accurately correct for these interferences, we doped double spiked DROsS with a range of 267 elements that are (i) typically found in silicate rocks (Si, Al, Mg, Fe, Ca, Ti, Ni and Cr), (ii) dominant in 268 iron meteorites (Fe, Ni), (iii) platinum group elements (Ru, Rh, Pd, Ir, and Pt), and (iv) trace elements 269 that have direct isobaric interferences on Os (Pt, W and Re). Test solutions were doped at levels ranging 270 from 0.01-10 % of the total Os concentration. Each element was present in x% of the Os abundance, 271 e.g. Si was present at 5% of the Os abundance. All measurements display values that are within 272 analytical uncertainty of undoped DROsS measurements (Fig. 3; Table 3). This illustrates that matrix 273 effects as well as isobaric interferences, even at extreme levels of up to 10 %, can be accurately 274 accounted for. The interference beam intensities monitored for standard solutions as well as geological 275 sample material processed and analysed in this study are all <0.01 % of the total Os concentration and 276 are, therefore, not expected to introduce data inaccuracies.

277

278 4.2 N-TIMS

279 4.2.1 Instrument set-up and parameters

280 For N-TIMS analyses, sample material was loaded onto Pt single filaments in concentrated HBr or in 3 281 M HCl when the sample solution was also measured by MC-ICP-MS. After the sample material was dried 282 down on the filament, ~0.5 µL of NaOH-Ba(OH)₂ activator was added to enhance ionisation. The amount 283 of Os loaded for this study varied between 2.3 and 45 ng natural Os. Where MC-ICP-MS provides fairly 284 consistent beam intensities for a certain set-up, the intensity that is obtained for a given amount of Os 285 is not directly correlated on N-TIMS, but also depends on the activator and sample loading on the 286 filament. A two-sequence static multi-collection routine was used in order to analyse masses 232 to 287 242 (Table 2). Masses 241 and 242 are collected to calculate the oxygen isotope compositions (see 288 section 4.2.2). A single analysis consists of 28 blocks, each comprising 10 cycles with 8.389 seconds 289 integration per cycle followed by 3 seconds idle time. Amplifier gain calibrations were performed at the 290 start of each day, although amplifiers were rotated throughout an analysis to cancel out amplifier gains. 291 Baseline measurements and peak centering, using masses 236 and 240, were carried out prior to every 292 run. The presence of interfering PtO₂⁻, ReO₃⁻, and WO₃⁻ were quantified by measuring masses 228, 230-293 233 by ion counter before and after each Faraday measurement for 4-6 cycles with a total analysis time of 110-160 seconds. During analysis, high purity oxygen (99.6 % purity, supplied by BOC) was bled into
 the source chamber with the pressure kept constant at ~2.5 x 10⁻⁷ mbar.

296

297 4.2.2 Data reduction N-TIMS

298 Osmium analyses by N-TIMS are performed using the tri-oxide ion (OsO₃⁻) and, therefore, corrections 299 for the effects of variable isobaric oxygen isotope interferences of the heavier oxygen isotopes (¹⁷O and ¹⁸O) are required. For example, correction for ${}^{188}Os^{16}O_2{}^{17}O^{-}$ on ${}^{189}Os^{16}O_3^{-}$ at mass 239. In this study, we 300 301 adopted the method of Luguet et al. (2008)³⁷ where the O isotope composition is determined in-run for 302 each integration so that any variation throughout an analysis can be accounted for. This is a more 303 accurate approach than using a fixed O isotope composition because the composition has been shown 304 to vary from sample to sample as well as throughout an individual run e.g. 37, 39, 51, 52. For comparison, we 305 have also applied the method recommended by Chatterjee and Lassiter (2015)³⁹ where the O isotopic 306 composition was measured before and after the main run using a separate routine. This routine 307 consisted of the second line as mentioned in Table 2 and data was collected for 10 cycles of 8.389 308 seconds integrations each. From the 20 cycles, the mean 241/238 and 242/238 values (2 se outlier 309 rejected) were taken as an approximation of the O isotopic composition of the specific run. The 310 collection of masses 241 and 242, which are free from the most abundant Os tri-oxide species $({}^{16}O_3)$, 311 allows the O isotope compositions to be determined by stripping masses 240, 241 and 242 of the minor tri-oxide interferences, such as ¹⁸⁹Os¹⁶O¹⁸O₂⁻ and ¹⁹⁰Os¹⁷O₃⁻ for mass 240. This will then provide the 312 signal intensity of only ${}^{192}Os^{16}O_{3}^{-}$, ${}^{192}Os^{16}O_{2}^{17}O^{-}$, and ${}^{192}Os^{16}O_{2}{}^{18}O^{-}$ on respective masses 240, 241, and 313 314 242. For this first step of oxygen corrections we used the O_{gas} isotope composition as determined by Luguet et al. (2008)³⁷. The ¹⁸O/¹⁶O was then calculated from the "stripped" 242/240 ratio and the 315 316 ¹⁷O/¹⁶O ratio could be determined from the "stripped" 241/240 ratio. However, as discussed in more 317 detail in section 4.2.2.1, in this study ¹⁷O/¹⁶O ratios were inferred using ¹⁸O/¹⁶O ratios. The oxygen 318 correction is described in detail in the Electronic Supplementary Information.

319 Due to the presence of isobaric Os oxide complexes, additional steps in the data reduction 320 compared to MC-ICP-MS were required. We adopted the following sequence of corrections; abundance 321 sensitivity (0.3 ppm³⁷), a preliminary oxide isotope composition correction, PtO_2^- and ReO_3^- interference 322 corrections, a second O isotope composition correction, and double spike deconvolution (see section 323 4.1.2). The first O interference correction is solely used to correct for O isobaric interferences on PtO2. Subsequently, the PtO2⁻ interference corrected ¹⁸⁴OsO3⁻ and ¹⁸⁶OsO3⁻ values were used in the second O 324 325 correction. For isobaric PtO₂⁻, WO₃⁻ and ReO₃⁻ interference corrections, the ion counter measurement 326 routines (before and after) were utilized, assuming a linear behaviour throughout the main run. Using 327 a linear extrapolation is a simplistic approach but can be justified by the minor variation in beam 328 intensities between the two SEM runs in combination with the minor influence the interference 329 corrections have on the Os isotopic ratios (see section 4.2.2.2). The abundance sensitivity was assumed 330 to be similar to that determined by Luguet et al. (2008)³⁷.

332 4.2.2.1 Effect of the oxygen isotopic composition

The accuracy of the deconvolved δ^{190} Os value as well as of the radiogenic isotopic ratios is dependent on the accurate determination of the O isotopic compositions (i.e. 17 O/ 16 O and 18 O/ 16 O) that are used for O interference corrections.

336 The average ¹⁸O/¹⁶O ratios of individual double spiked DROsS analyses range between 337 0.002007 and 0.002050, a variation of 21 ‰, for analyses with a total Os beam intensity of >1 V. This 338 variation is comparable with that reported by previous Os studies by N-TIMS (e.g. Liu et al., $1998^{53} - 30$ 339 ‰; Luguet et al., 2008³⁷ – 14 ‰; Chatterjee and Lassiter, 2015³⁹ – 31 ‰; Chu et al., 2015⁵⁴ – 22 ‰). If 340 the ${}^{18}O/{}^{16}O$ ratio varies by 21 ‰, and the ${}^{17}O/{}^{16}O$ relationship is changed accordingly along a TFL, then 341 the absolute δ^{190} Os value will shift by ±0.11 ‰. For the ¹⁸⁷Os/¹⁸⁸Os and ¹⁸⁶Os/¹⁸⁸Os ratios the resulting 342 absolute shift is about ± 20 ppm and ± 25 ppm, respectively. It is, therefore, of high importance to determine the specific oxygen isotopic composition of an analysis. When ^{total}Os beam intensities are on 343 344 average >1 V (>1.3 mV on mass 242), the difference in $^{18}O/^{16}O$ ratios between the first and last 20 cycles 345 of an individual run ranges, apart from two samples at -12 ‰, between +5 ‰ and -7 ‰, and is on 346 average -1.1 ‰. This is similar to the variation observed by Chatterjee and Lassiter (2015)³⁹. Larger 347 variations are observed for measurements made at lower beam intensities and the variability appears 348 to be more substantial when samples are run over a longer time period (Fig. 4). If the $^{18}O/^{16}O$ ratio is changed by ±7 ‰, the absolute composition of δ^{190} Os shifts by ±0.025 ‰. The ¹⁸⁷Os/¹⁸⁸Os and 349 350 186 Os/ 188 Os ratios are shifted by ±11 ppm and ±26 ppm, respectively. We, therefore, suggest 351 determining the oxygen isotopic composition for each individual cycle, as this is the only way by which 352 small time scale variations throughout the run can be monitored and controlled.

353 Run-specific ¹⁷O/¹⁶O ratios for each individual cycle can either be measured directly, using 354 mass 241, or can be calculated based on the ¹⁸O/¹⁶O ratios. We will first consider the measured ¹⁷O/¹⁶O 355 ratios. For DROsS analyses, the average measured ¹⁷O/¹⁶O ratio was found to display a large variation, of 121 ‰, with values ranging from 0.000369 to 0.000419. When only considering analyses with total 356 357 beam intensities of >1 V the $^{17}O/^{16}O$ ratios vary between 0.000369 and 0.000395, 66 %. Analyses with 358 an anomalously high or low ${}^{17}O/{}^{16}O$ ratio yield inaccurate $\delta^{190}Os$ values as well as ${}^{187}Os/{}^{188}Os$ and 359 ¹⁸⁶Os/¹⁸⁸Os ratios, and show a positive correlation with ¹⁷O/¹⁶O (R² of 0.78, 0.78 and 0.76 respectively; 360 Fig. 5a-c). Anomalous values are only obtained for analyses performed at low beam intensities (Fig. 5df). In contrast, the measured ¹⁸O/¹⁶O ratios do not show any co-variation with Os isotopic compositions 361 362 (R²=0.28) suggesting that the inaccuracy on the Os isotopic compositions is dominantly introduced by 363 the ¹⁷O/¹⁶O ratio. Mass 241 has been measured at very low beam intensities (<5 mV) that are, in general, 5.2 times lower than for mass 242. The determination of $^{17}O/^{16}O$ is, therefore, more sensitive to 364 365 inaccuracies introduced by small variations in the instrument baseline occurring during an analysis. 366 Including more and longer baselines with longer integration times during an analysis would enhance 367 the accuracy of the baseline and potentially improve the measurement of the ¹⁷O/¹⁶O ratio. 368 Incorporation of $10^{12} \Omega$ or $10^{13} \Omega$ resistors to collect mass 241 and 242 would also be advantageous, 369 although these were not explored in this study.

370 The inaccuracy introduced by the measured ${}^{17}O/{}^{16}O$ ratio can be eliminated by calculating the 371 $^{17}\text{O}/^{16}\text{O}$ from the measured $^{18}\text{O}/^{16}\text{O}$ ratio. We will refer to this as the "calculated" $^{17}\text{O}/^{16}\text{O}$ ratio. 372 Recently, a similar approach has been used for N-TIMS isotope measurements of Mo⁵², Ru⁵⁵, and W⁵⁶. 373 In this study, we assume that the oxygen isotopic compositions vary as a result of equilibrium massdependent isotopic fractionation along a "Terrestrial Fractionation Line" (TFL⁵⁷). For natural samples 374 375 the relationship between δ^{17} O and δ^{18} O, when considered in logarithmic space, typically varies between 376 0.524-0.528 ^{e.g.58-60}. We have taken an average of the published values, 0.526, which translates to an approximately linear slope of 0.0994 x + 0.000183, where x is ${}^{18}O/{}^{16}O$, over a ${}^{18}O/{}^{16}O$ interval of 377 378 0.002007 to 0.002055, the range in ${}^{18}O/{}^{16}O$ as observed for analyses in this study. Figure 6 shows that 379 the measured oxygen compositions of double spiked DROsS analyses with a total Os beam intensity of 380 >2 V closely plot along this line. Details of calculation $^{17}O/^{16}O$ from the measured $^{18}O/^{16}O$ are provided 381 in the supplementary information.

382 The resulting calculated $^{17}O/^{16}O$ ratios display a much smaller variation of 20 % with values 383 ranging between 0.000383 and 0.000390, and is 10 ‰ when only considering analyses of >1 V. No residual correlation between calculated ¹⁷O/¹⁶O ratios and Os isotopic compositions is observed 384 385 anymore (Fig. 5a-c). However, with decreasing beam intensities the δ^{190} Os values display a preferential 386 drift towards heavier values (up to 0.11‰) instead of showing a normal distribution (Fig. 5d). The 387 ¹⁸⁷Os/¹⁸⁸Os and ¹⁸⁶Os/¹⁸⁸Os, on the other hand, drift preferentially towards lower values although higher 388 values are observed as well (Fig. 5e,f). When the calculated instead of the measured $1^{7}O/1^{6}O$ 389 composition is used, the long-term reproducibility of δ^{190} Os is improved from 0.106 ‰ to 0.046 ‰ (2 390 sd; n = 94), and from 0.059 ‰ to 0.029 ‰ for analyses of >1 V (n = 83). For the remainder of this paper, 391 we will only consider analyses with a totalOs beam intensity of >1 V, equivalent to, in general, >1.3 mV 392 on mass 242 and >0.35 V on mass 236. The total Os signal intensity obtained for a certain quantity of 393 Os loaded is highly variable, but is typically >1 V for loads of \ge 2.3 ng natural Os.

394

395 4.2.2.2 Effect of polyatomic interferences

396 The interference intensities monitored on masses 228, and 230-233 in this study are comparable to 397 those reported by Luguet et al. (2008)³⁷. Beam intensities on mass 228 (predominantly ¹⁹⁶Pt¹⁶O₂⁻) 398 ranged from 12,000 – 1,400,000 cps (counts per second) which relates to a potential ¹⁹⁶Pt¹⁸O₂-399 interference of 0.05-6 cps on mass 232 (predominantly ¹⁸⁴Os¹⁶O₃⁻). For mass 230 (mainly ¹⁹⁸Pt¹⁶O₂⁻) 400 intensities ranged between 4,000-1,050,000 cps which results in respective interferences of 16-4,300 cps and 0.02-4.4 cps on masses 232 (¹⁸⁴Os¹⁶O₃⁻) and 234 (¹⁸⁶Os¹⁶O₃⁻). The PtO₂⁻ interference on mass 234 401 402 is negligible, whereas the combined PtO_2^- interferences on mass 232 reached a maximum of 4,306 cps. This represents 2.5 % of the typical total signal at mass 232 which can lower the ¹⁸⁴Os/¹⁸⁸Os substantially 403

404 (~30,000 ppm). In this study, mass 234 was only monitored to correct for polyatomic oxygen 405 interferences on the other Os isotopic ratios. Lowering the ¹⁸⁴Os/¹⁸⁸Os by 30,000 ppm results in a minor shift of -2 ppm on the ¹⁸⁶Os/¹⁸⁸Os and has no noticeable effect on the other Os isotopic ratios. The 406 407 average difference between the two SEM runs, before and after the main run, was 250,000 cps for mass 408 228 and 135,000 cps for mass 230. Potential within-run variations of this magnitude have no 409 measurable effect on the isotopic ratios. On mass 231, ¹⁸³WO₃⁻, we obtained intensities of 6-1,000 cps, which produce insignificant interferences on ¹⁸⁶Os¹⁶Os^{- 37, 61}. Mass 233, ¹⁸⁵ReO₃-, displayed intensities 410 411 between 3-460 cps but were generally below 300 cps and on average 73 cps. Translated to ¹⁸⁷ReO₃⁻ this 412 means <502 cps or 8 μ V that interfered on the ¹⁸⁷OsO₃. Typically, this quantity equates to a lowering of the ¹⁸⁷Os/¹⁸⁸Os by <100 ppm and on average 38 ppm, which is close to the 2 se of our measurements 413 (40 ppm) but much smaller than the long-term reproducibility (268 ppm; 2 sd). Between the pre- and 414 415 post SEM runs, the beam intensities on mass 233 generally varied by 33 cps which equates to a shift of ~18 ppm on the 187 Os/ 188 Os. To summarize, polyatomic interferences from PtO₂, WO₃, and ReO₃ 416 monitored in this study had no noticeable effect on the stable Os and ¹⁸⁶Os/¹⁸⁸Os isotopic composition, 417 418 and were minor for ¹⁸⁷Os/¹⁸⁸Os ratios.

419

420 5. RESULTS AND DISCUSSION

Results are presented in Tables 4 and 5, and shown in Figures 7-13. All internal errors are quoted as 2
standard error (2 se), whereas short term (i.e. single session) and long-term (i.e. multiple sessions)
reproducibilities are given as 2 standard deviations (2 sd).

424

425 **5.1** Internal precision, external reproducibility and accuracy

426 **5.1.1 Stable osmium isotope ratios by MC-ICP-MS**

427 The internal precision (2 se; $n \le 80$ cycles, depending on 2 se outlier rejection) on a single δ^{190} Os MC-ICP-MS analysis is typically between 0.01-0.02 ‰ when ¹⁸⁸Os beam intensities range between 7-4 V 428 429 (~18-11 V totalOs; Fig. 7). The amount of natural Os consumed is ~200 ng. The observed precision is in 430 good agreement with the theoretical error as calculated in section 2.1 (Fig. 7) suggesting that the model 431 considers all the errors that should be accounted for. Small deviations from the calculated error could 432 have arisen from, for example, variation in the spike – sample proportions, or the number of cycles 433 included. The model considers 80 cycles whereas this can be less for MC-ICP-MS if outliers are rejected. 434 This also explains why the error model for N-TIMS displays lower errors for a given intensity, as 280 435 cycles were considered for N-TIMS analyses.

436 Repeated analyses of reference standard solution DROsS, obtained during multiple analytical 437 sessions over a time period of ~22 months, show an external reproducibility on δ^{190} Os of 0.016 ‰ (2 438 sd; *n* = 91; Fig. 8; Table 4). The short-term reproducibility of a single analytical session when consuming 439 ~200 ng natural Os at total beam intensities of 11-18 V is 0.014-0.029 % (2 sd; n = 2-10). A similar 440 precision and reproducibility is obtained for in-house Os reference solutions ROMIL, SpecPure and 441 OsCaR (Fig. 7 and 9; Table 4). As this study is the first to present stable Os isotope compositions, the 442 accuracy cannot be assessed by measurement of pre-calibrated reference materials. Instead we have 443 performed standard-sample bracketing measurements by MC-ICP-MS using natural (i.e. non DS) 444 reference solutions. Comparison of this technique with the DS method shows that a similar offset 445 between DROsS and the other Os reference solutions is obtained (Fig. 9; Table 4) which is in support of 446 the relative accuracy of our method.

447

448 5.1.2 Stable osmium isotope ratios by N-TIMS

449 For a single δ^{190} Os analysis by N-TIMS the internal precision (2 se; n \leq 280 cycles, depending on 2 se 450 outlier rejection) ranges between 0.010-0.040 % when ¹⁸⁸OsO₃⁻ beam intensities vary from 6-0.43 V 451 (~18 to 1 V totOs beam; Fig. 7). These intensities were obtained for load sizes between 45 and 2.3 ng 452 natural Os. Repeated analyses of double spiked reference solution DROsS, over approximately 22 453 months, yield an external reproducibility on δ^{190} Os of 0.029 ‰ (2 sd; n = 83; Fig. 8; Table 4). The majority 454 of these analyses (n = 44) comprised a load of 10-20 ng natural Os, 20 analyses had a load of >20 ng, 455 and 19 analyses were performed with less than 10 ng Os. That DROsS yields less precise and less 456 reproducible results for N-TIMS than MC-ICP-MS is, predominantly, a consequence of the lower beam 457 intensities at which the analyses were performed. Analyses of in-house pure Os solutions ROMIL, SpecPure and OsCaR show a similar precision and reproducibility (Fig. 7 and 9; Table 4). That ROMIL 458 459 and SpecPure display a similar offset relative to DROsS for N-TIMS as for MC-ICP-MS analyses provides 460 support to the accuracy of our method.

461

462 5.1.3 Radiogenic isotope ratios by MC-ICP-MS

The precision that can be obtained for 187 Os/ 188 Os and 186 Os/ 188 Os isotopic ratios is <100 ppm (2 se; n \leq 463 280 cycles, depending on 2 se outlier rejection) when ¹⁸⁷Os and ¹⁸⁶Os average beam intensities are 464 465 higher than ~0.06 V over the ~11 minutes of the analysis. When average beam intensities of >0.18 V 466 are achieved this improves to ~40 ppm (Fig. 7). This is comparable with the precisions reported by Nowell et al. (2008; <40 ppm at high signal intensities)³⁶ and corresponds with the errors predicted by 467 468 the model described in section 2.1 (30-40 ppm at 0.22 V; Fig. 7). Repeated analyses of double spiked 469 reference standard DROsS yields a reproducibility of 123 and 138 ppm for ¹⁸⁷Os/¹⁸⁸Os and ¹⁸⁶Os/¹⁸⁸Os, 470 respectively (2 sd, n = 91). This is higher than that reported by Nowell et al. (2008)³⁶ for DROsS (19 ppm) 471 but is comparable with the long-term reproducibilities reported for reference materials UMd (220 ppm for ¹⁸⁷Os/¹⁸⁸Os and 108 ppm for ¹⁸⁶Os/¹⁸⁸Os) and DTM (149 ppm for ¹⁸⁷Os/¹⁸⁸Os and 67 ppm for 472 ¹⁸⁶Os/¹⁸⁸Os). 473

474 For radiogenic isotopic ratios we can use previous studies to assess the accuracy of our 475 method. In addition, we have compared double spiked data with that from unspiked runs, and used 476 DROsS for un-spiked sample-standard bracketing measurements. We have applied these methods to 477 DROsS as well as to pure Os solutions ROMIL, SpecPure, and OsCaR. Taken together, the various samples encompass a significant range in ¹⁸⁷Os/¹⁸⁸Os isotopic ratios over which to test our analytical methods. 478 479 Table 4 shows that all values obtained in this study, using the various methods, are within analytical 480 uncertainty with one another as well as with previously published values of un-spiked DROSS 481 measurements.36,37

482

483 **5.1.4 Radiogenic isotope ratios by N-TIMS**

For N-TIMS analyses the internal precision on both 187 Os/ 188 Os and 186 Os/ 188 Os is <100 ppm (2 se; n \leq 484 485 280 cycles, depending on 2 se outlier rejection) for analyses with average 187 OsO₃⁻ and 186 OsO₃⁻ beam 486 intensities >0.03 V over the 280 cycles of analysis, and improves to <40 ppm at high signal intensities 487 (>0.18 V; Fig. 7). This is slightly greater than the precisions reported by Luguet et al. $(2008^{37}; <30 \text{ ppm})$ 488 at >0.08 V). Repeated analyses of reference standard DROsS yield a reproducibility of 268 ppm and 361 489 ppm for 187 Os/ 188 Os and 186 Os/ 188 Os ratios, respectively (2 sd, n = 83) for loads varying between 2.3 and 490 45 ng natural Os. This is significantly higher than the values reported by Luguet et al. (2008³⁷; 26 and 48 491 ppm, respectively). Where Luguet et al. (2008)³⁷ have only included analyses with ¹⁸⁷OsO₃⁻ and ¹⁸⁶OsO₃⁻ 492 beam intensities of 80 mV or above we have taken analyses with intensities down to 20 mV into account. When excluding the analyses with beam intensities <80 mV on ¹⁸⁷OsO₃- and ¹⁸⁶OsO₃- we obtain 493 494 a long-term reproducibility of 90 and 89 ppm, respectively (2 sd; n = 22). These slightly higher analytical 495 uncertainties are related to error propagation inherent in double-spike deconvolution. That errors are 496 higher than predicted by the model is, predominantly, because the error on the oxygen composition 497 was not incorporated in the model, whereas it has been propagated for the actual analyses.

To assess the accuracy, we have compared the radiogenic values obtained by our DS method, with those determined for unspiked runs, and with previous studies (Table 4). All values are within analytical uncertainty of one another which demonstrates that our method is able to obtain accurate ¹⁸⁷Os/¹⁸⁸Os and ¹⁸⁶Os/¹⁸⁸Os isotopic ratios by N-TIMS. This provides a valuable quality control on the DS calculations and allows the acquisition of both the stable and radiogenic isotope composition within a single analytical run.

504

505 5.2. Geological materials

506 During the course of this study, we have measured four international geological reference materials. 507 The selected materials cover a range of matrices; (1) UB-N⁴⁰, a serpentinised and fertile lherzolite from 508 the Voges in France; (2) GP-13^{41, 62}, a fertile lherzolite from the Beni Bousera massif in Morocco; (3) 509 CHR-Bkg ^{42, 43} and (4) CHR-Pt+^{42, 63}, which are both chromitites from the Shetland ophiolite in Scotland, 510 UK. Furthermore, we have analysed the ordinary H-chondrite Zag. Although this is not an international 511 reference material, sufficient sample material was available to perform several replicate analyses 512 allowing the exploration of another type of sample matrix. Stable and radiogenic Os isotope 513 compositions, together with Os abundances, are given in Table 5 and shown in Fig. 10-13.

514

515 **5.2.1** Radiogenic osmium isotope compositions for geological materials

516 We have shown that double spiked DROsS analyses display ¹⁸⁷Os/¹⁸⁸Os and ¹⁸⁶Os/¹⁸⁸Os ratios consistent 517 with previous published studies, and that ROMIL and SpecPure yield similar values for various methods 518 (i.e. non-DS, DS, standard sample bracketing). This means that although the main aim of this method is 519 to obtain high precision stable isotope data it also has the potential to provide radiogenic isotopic ratios 520 within the same analyses. This is particularly beneficial when dealing with limited amounts of available 521 sample material that only allows a single analysis. Furthermore, it significantly reduces processing and 522 measurement time.

523 Apart from three UB-N analyses, all data is obtained at average ¹⁸⁷OsO3⁻ and ¹⁸⁶OsO3⁻ beam 524 intensities of \geq 0.01 V for N-TIMS (^{total}Os \geq 1 V) analyses and ¹⁸⁷Os and ¹⁸⁶Os \geq 0.03 V for MC-ICP-MS (totalOs \geq 1 V). The influence of the ${}^{17}O/{}^{16}O$ composition and OPZ on the accuracy of the isotopic 525 526 compositions is shown to be negligible at these intensities. Geological materials analysed in this study display ¹⁸⁷Os/¹⁸⁸Os values that are consistent with the range of values obtained by previous studies (Fig. 527 528 10 and 13). However, a significant degree of irreproducibility (up to 7,400 ppm) can be observed which 529 will be discussed below. The variability observed for ¹⁸⁶Os/¹⁸⁸Os ratios is much smaller, 251-537 ppm, 530 and comparable with the long-term reproducibility obtained for N-TIMS DROsS analyses (361 ppm). The 531 186 Os/ 188 Os ratios of all samples range from 0.119746 ± 42 to 0.119856 ± 47, which is in close 532 approximation of the upper mantle value (0.119837±5)⁶⁴ but consistently lower (Fig. 11). Assessment of the accuracy of our ¹⁸⁶Os/¹⁸⁸Os values by direct comparison with previously published values for these 533 534 samples is not possible as they have not been measured for ¹⁸⁶Os/¹⁸⁸Os before. At this point, the reason 535 for the lower values is unclear and further investigation is required.

536

537 **5.2.2** Stable osmium isotope compositions for geological materials

The internal precision and external reproducibility on δ^{190} Os are similar to that obtained for pure Os isotope solutions. One exception is the reproducibility reported by CHR-Pt+. This will be discussed in more detail in the next section. Similar to Os standard solutions, replicate analyses of chondrite ZAG show similar values for MC-ICP-MS and N-TIMS analyses, providing further support for the accuracy of our method on a real sample matrix.

543 All reference materials are derived from Earth's mantle. The materials show no resolvable 544 variation at the 95 % c.i. level and provide an average δ^{190} Os value of 0.130 ± 0.032 ‰ (2 sd; *n* = 4). The geological reference material CHR-Bkg displays the lightest composition of δ^{190} Os = 0.124 ± 0.020 ‰ (2 sd; *n* = 7) and chromitite CHR-Pt+ the heaviest at 0.162 ± 0.051 (2 sd; *n* = 4). Ordinary chondrite Zag yields a δ^{190} Os value of +0.123 ± 0.018 ‰ (2 sd; *n* = 9) which is indistinguishable, within uncertainty, from the mantle derived samples.

549 Interestingly, DROsS is the only material that displays a different, isotopically lighter, 550 composition relative to the other materials measured in this study. This offset could (a) be a product of 551 the chemical extraction of Os from the source material, or (b) reflect the original composition of the 552 material the Os was sourced from.

553

554 **5.2.3 Effect of sample digestion**

555 In mantle rocks, osmium is concentrated in refractory accessory phases that are heterogeneously 556 distributed throughout the rock (the "nugget" effect). This has led to difficulties in repeating results when considering Os abundances and ¹⁸⁷Os/¹⁸⁸Os ratios. Over the past two decades, various digestion 557 methods have been assessed in order to resolve this problem ^{e.g. 40, 44, 65, 66}. Acid attack digestions using 558 559 sealed glass Carius tubes (CT^{23, 45}) and the high-pressure asher system (HPA⁴⁴) are most frequently used. This is because they are considered to be most efficient in attacking the highly resistant phases that are 560 561 likely to contain appreciable Os (such as Cr-spinel and platinum-group minerals). In this study, we have 562 applied both techniques to various geological materials. A limitation of both techniques is the maximum 563 amount of sample material that can be digested effectively (≤ 2 g). In the case of geological samples 564 with low Os abundances (e.g. most mantle material like UB-N, and nearly all crustal material) more than 565 2 g of material is needed to obtain high precision data. To circumvent this problem, we have digested 566 several sample aliquots of UB-N and GP-13, and combined them prior to solvent extraction. As for the 567 individual processed samples, each aliquot was spiked before digestion. As the proportion of spike is 568 one of the unknowns calculated during the resolution of the DS equation system, it should be noted 569 that reaching a 100 % yield across all aliquots during sample processing is not required to generate 570 accurate isotope ratios and concentration determinations, providing that spike-sample equilibration 571 occurs before loss of any Os during sample processing.

572 Apart from CHR-Pt+, replicate digestion of individual samples shows no detectable stable 573 isotope variation and no systematic difference between CT or HPA digestions. For the combined 574 aliquots of UB-N, relative to the single digestions, we do not observe any variation for stable Os isotopes either (Fig. 12; Table 5). The two analyses of combined GP-13 aliquots show consistent stable isotope 575 compositions. By contrast, statistically resolvable variations in ¹⁸⁷Os/¹⁸⁸Os ratio and Os abundance were 576 577 observed in all the samples tested. In the case of UB-N and CHR-Bkg, ¹⁸⁷Os/¹⁸⁸Os is positively correlated 578 with 1/[Os] (Fig. 13). This co-variation was observed previously, in a more extensive study of UB-N, by 579 Meisel et al. (2003)⁴⁰. The CHR-Bkg analyses reported within that study display a more limited variation 580 and no significant co-variation between Os abundance and radiogenic isotopic composition was

discussed. If we incorporate the CHR-Bkg data of Meisel et al. (2003)⁴⁰ they fall within the trend 581 582 obtained in this study. The observed variation in radiogenic isotopes can be explained by incomplete 583 digestion and/or sample heterogeneity. In both cases this indicates that within the digested material 584 different phases, or phases of different generations, possess distinct radiogenic isotopic composition. 585 The absence of stable Os isotope variation, for these particular aliquots, implies that there is no 586 significant difference in stable isotope composition of the phases in these samples. As such, the type of digestion technique, or the combination of individual digestions, appears to have no influence on the 587 588 stable isotope composition for these particular samples.

589 For CHR-Pt+ we observe a larger range in stable Os isotope compositions (δ^{190} Os = +0.129 -590 0.198 ‰), which is reflected in the relatively poor reproducibility (0.051‰). The stable isotope data 591 appear to co-vary with Os abundance, with the heavier values corresponding to higher Os abundance 592 of 17.5 % (Fig. 12). As discussed before, variance in isotopic composition correlating with Os abundance 593 may indicate that phases within the whole rock contain different stable isotopic signatures. Our 594 preliminary data suggests that phases within chromitite CHR-Pt+ possess different stable Os isotopic 595 compositions. Previous studies that focused on the origin of chromitites derived from the Cliff deposit 596 in the Shetland Ophiolite concluded that initial PGE concentrations were caused by magmatic processes 597 followed by a hydrothermal overprint which locally remobilized and re-concentrated the PGE.⁶⁷ This 598 secondary process might be the source of Os stable isotope fractionation, but further investigation is 599 required.

600

601 6. CONCLUSIONS

602 We have developed a method for high-precision measurement of stable Os isotope compositions by 603 both plasma source (MC-ICP-MS) and thermal ionisation mass spectrometry (N-TIMS). The method utilizes a ¹⁸⁸Os-¹⁹⁰Os double spike that is added to the sample prior to digestion with a spike-sample 604 605 proportion of 0.55:0.45. We show that the technique is robust when dealing with matrix effects and 606 interference corrections on MC-ICP-MS, even for levels up to 10 % of the total Os concentration, and 607 that the memory effect for MC-ICP-MS analyses is negligible. Analyses performed by N-TIMS show that 608 the oxygen isotopic composition exerts a major control on the accuracy of the isotopic ratios. It is 609 suggested that the oxygen isotopic composition is obtained for every specific run, by measuring the 610 $^{18}\text{O}/^{16}\text{O}$ for each individual cycle, and calculating the $^{17}\text{O}/^{16}\text{O}$ ratio from the $^{18}\text{O}/^{16}\text{O}$ ratio, especially 611 when the total Os signal intensity drops below 2 V.

The internal precision (2 se) on the δ^{190} Os measurement (permil difference of the 190 Os/ 188 Os ratio relative to reference standard DROsS) of a single analysis is 0.010-0.025 ‰ for MC-ICP-MS (~200 ng natural Os; n = 80 cycles) and 0.010-0.030 ‰ for N-TIMS (2.3-45 ng natural Os; n = 280 cycles). The long-term reproducibility of reference material DROsS is 0.016 ‰ (2 sd; n = 91) and 0.029 ‰ (2 sd; n =616 83) for MC-ICP-MS and N-TIMS, respectively. 617 The method is shown to be capable of obtaining Os stable isotope compositions of terrestrial 618 and extra-terrestrial materials with a high precision and reproducibility. The first data obtained for 619 geological materials are all within analytical uncertainty of one another, despite the measurement of 620 samples reflecting a range of source regions, chemical compositions, geological history, and radiogenic 621 isotope ratios. Minor variations between different digestions of chromitite CHR-Pt+ hints at Os stable 622 isotope fractionation as a result of Os remobilization due to hydrothermal alteration. Overall, our 623 preliminary δ^{190} Os value for the Earth's upper mantle is +0.130 ± 0.032 ‰ (2 sd; n = 4), which is 624 indistinguishable from a value of +0.123 \pm 0.018 % (2 sd; n = 9) obtained for the ordinary H-chondrite 625 Zag.

626

627 Acknowledgements

We acknowledge financial support from an ERC starting Grant awarded to H. Williams (Habitable Planet
306655), which funded most of this analytical work, a Marie Curie COFUND International Junior
Fellowship granted to M.-A. Millet, and a Durham University Scholarship awarded to J. Nanne. We thank
C. Cloquet for providing us UB-N and A. Poirier for OsCaR. D. Selby, E. Dempsey, A. du Vivier, and A.
Sproson are thanked for assistance in getting to grips with CT digestion and sample preparation, and C.
Ottley for maintenance of the HPA.

634

636 FIGURE CAPTIONS

Figure 1 Results for an error (2 se) simulation on the stable isotope composition in % amu⁻¹ in case of a ¹⁸⁸Os-¹⁹⁰Os double spike. The red dot indicates that minimal error propagation of ~0.010 % amu⁻¹ (2 se) is obtained when using a 60 % ¹⁸⁸Os – 40 % ¹⁹⁰Os spike that is mixed with a sample in relative proportions of 55 % and 45 %. Note that errors are minimal (<0.020 % amu⁻¹) over a large range of sample-spike mixtures (15 – 81 % sample fraction). The error model is based on the method of Millet and Dauphas (2014)³³ using the following parameters: 6 V; 80 x 8.3 sec integrations; 10¹¹ Ω collectors; T= 290 K.

Figure 2 Results of double spiked DROsS analyses measured using MC-ICP-MS (closed symbols) at variable beam intensities to assess the effect of beam intensity and the on-peak zero on the accuracy of the δ^{190} Os and 187 Os/ 188 Os composition. Data show that all measurements display $\delta^{190/188}$ Os and 187 Os/ 188 Os compositions within analytical uncertainty of the long-term reproducibility as determined for >10 V analyses. The black dotted line indicates the average DROsS value, as obtained from the >10 V analyses, with the ±2 sd represented by the grey band.

Figure 3 Results of doping tests to assess the robustness of the method against residual matrix effects on the MC-ICP-MS. Data show that all measurements display $\delta^{190/188}$ Os compositions within analytical uncertainty of undoped DROsS analyses for residual element abundances up to 10 % of the total Os concentration. The percentages indicate the presence of each named element at x% of the Os abundance, e.g. Si was present at 5% of the Os abundance. The black dotted line indicates the average DROsS value with the ±2 sd represented by the grey shaded band as obtained in this study. Lith = lithophile elements (Si, Al, Mg, Fe, Ca, Ti, Ni, and Cr); PGE = Platinum Group Elements (Ru, Rh, Pd, Ir, and Pt).

Figure 4 The difference in measured ¹⁸O/¹⁶O ratios between the first and the last 20 cycles of double spiked DROSS analyses (green circles) show an average variability of -1.1 % (green dotted line). For clarity, two values with a deviation of -12 %, at 1.4 V and 1.7 V, are excluded. The variability is shown to increase with decreasing beam intensity and seems to be larger when samples are run twice for 280 cycles. Open symbols represent analyses where the oxygen composition has been measured before and after the main run.

Figure 5 Residual correlation for double spiked DROsS analyses by N-TIMS between the measured ¹⁷O/¹⁶O (grey circles) and Os isotopic compositions; (a) δ^{190/188}Os, (b) ¹⁸⁷Os/¹⁸⁸Os, and (c) ¹⁸⁶Os/¹⁸⁸Os. No residual correlation is observed when the ¹⁷O/¹⁶O is calculated from the measured ¹⁸O/¹⁶O ratio (green circles). The offset from the expected DROsS value is seen to increase with decreasing beam intensities (d-f), with a reduced offset when using calculated ¹⁷O/¹⁶O ratios. With decreasing signal intensity, (d) δ^{190/188}Os values seem to slightly drift towards heavier values, and (e) ¹⁸⁷Os/¹⁸⁸Os and (f) ¹⁸⁶Os/¹⁸⁸Os towards lower values. The black dashed line represents the expected DROsS value.

667 Figure 6 Measured ¹⁸O/¹⁶O vs. ¹⁷O/¹⁶O ratios for DROsS analyses run at a total Os beam intensity of >2 V closely 668 follow the Terrestrial Fractionation Line (TFL). Analyses are divided based on the beam intensity on mass 242 (<5 669 mV; 5-10 mV; >10 mV). The TFL was calculated using a slope of 0.526 which translates to an roughly linear slope of 670 0.0994 x + 0.000183, where x is ${}^{18}\text{O}/{}^{16}\text{O}$, over a ${}^{18}\text{O}/{}^{16}\text{O}$ interval of 0.002007 to 0.002055, the range in ${}^{18}\text{O}/{}^{16}\text{O}$ 671 obtained in this study. Also shown are oxygen isotope ratios from N-TIMS measurements in previous studies: Nagai 672 and Yokoyama (2016)⁵¹; Worsham et al. (2016)⁵²; Luguet et al. (2008)³⁷; Griselin et al. (2001)⁶⁸; Chavagnac (1998)⁶⁹; 673 Thirlwall (1991)⁷⁰; Reisberg and Zindler (1986)⁷¹; Nyquist in Wasserburg et al. (1981); Wasserburg et al. (1981)⁷²; 674 Nier (1950)73.

675Figure 7 The analytical precision (2 se) on (a) $\delta^{190/188}$ Os and (b) 187 Os/ 188 Os plotted against the average (a) 188 OsO3⁻676and (b) 187 OsO3⁻ beam intensity (V) for analyses of reference material DROsS, and other pure Os solutions by MC-677ICP-MS or N-TIMS. The dashed (MC-ICP-MS) and dotted (N-TIMS) lines in (a) indicate the modelled error calculated678using the method of Millet and Dauphas (2014)³³ using the following parameters: 80 x 8.3 seconds integrations for679MC-ICP-MS and 280 x 8.3 seconds for N-TIMS; 10^{11} Ω collectors; T= 290 K. Note that the lower theoretical error for680N-TIMS results from the larger amount of cycles incorporated.

Figure 8 Repeated analyses of reference material DROSS on both MC-ICP-MS (squares) and N-TIMS (circles)
 obtained during multiple analytical sessions over a time period of 22 months. Analyses have been organized on
 beam intensity, note that N-TIMS analyses are obtained at lower beam intensities. Sample size is approximately

200 ng natural Os for MC-ICP-MS and ranges between 2.3 and 45 ng natural Os for N-TIMS measurements, with
 totalOs beam intensities of >4 V for MC-ICP-MS analyses and >1 V for analyses by N-TIMS. Error bars quote the 2 se
 error on the individual analysis, the grey band represents the ±2 sd of all analyses. Symbols in yellow correspond
 to the annotated beam intensity.

688Figure 9 The $\delta^{190/188}$ Os value obtained for repeat analyses of pure Os solutions ROMIL (triangle), SpecPure (circle)689and OsCaR (diamond) by both MC-ICP-MS (closed symbols) and N-TIMS (open symbols) obtained during the course690of this study. The striped, filled symbols indicate sample-standard bracketing data. All Os solutions show a similar691reproducibility as obtained for reference standard DROSS. Values obtained for ROMIL and SpecPure on MC-ICP-MS692and N-TIMS show values within analytical uncertainty supporting the accuracy of our method. The grey bands693indicate the reproducibility (±2 sd) obtained by MC-ICP-MS and N-TIMS. Error bars represent the 2 se error of an694individual analysis.

Figure 10 The ¹⁸⁷Os/¹⁸⁸Os isotope ratios of geological materials UB-N, GP-13, CHR-Bkg, CHR-Pt+ and Zag analysed
 by MC-ICP-MS (closed symbols) or N-TIMS (open symbols) display values consistent with previous studies (grey
 bars). Literature data for maximum and minimum values: UB-N – ref. 40, 74; GP-13 – ref. 75, 76; CHR-Bkg – ref 40;
 CHR-Pt+ - no previous published data; Zag – ref 3, 4.

Figure 11 The ¹⁸⁶Os/¹⁸⁸Os ratios of geological materials UB-N, GP-13, CHR-Bkg, CHR-Pt+ and Zag analysed by MC ICP-MS (closed symbols) or N-TIMS (open symbols). Most samples plot below the upper mantle value of
 0.119837±5 (2 sd; Ireland et al., 2011⁶⁴). CT = Carius tube digestions, all other sample aliquots are processed using
 the high-pressure asher system. * = an analysis for which digestions were combined prior to chemical Os extraction.
 Error bars indicate the 2 se on the individual analysis.

704Figure 12 Osmium stable isotope compositions ($\delta^{190/188}$ Os) of geological materials UB-N, GP-13, CHR-Bkg, CHR-Pt+705and Zag analysed by MC-ICP-MS (closed symbols) or N-TIMS (open symbols). Symbols are ordered from high to low706Os concentration. Apart from Zag, which is an ordinary chondrite, all samples are reference materials sourced by707the Earth's mantle. Symbols as in Fig. 11.

Figure 13 Correlation between ¹⁸⁷Os/¹⁸⁸Os ratio and 1/[Os] for reference materials UB-N (circles) and CHR-Bkg (triangles), as obtained for independently processed sample aliquots in this study (open symbols). Included are data as obtained by Meisel et al. (2003⁴⁰; filled symbols).

712 TABLE CAPTIONS

713 Table 1 Osmium isotopic ratios of reference standard DROsS, as determined by Nowell et al. (2008)³⁶ normalized 714 to ¹⁸⁹Os/¹⁸⁸Os to correct for mass bias, and of the ¹⁸⁸Os-¹⁹⁰Os double spike (DS) as used in this study. The quoted 715 error on DROsS represents the 2 sd on 21 analyses, and the error on the DS is the relative 2 se on the analysis used 716 to obtain the DS composition.

717

721

Table 2 Faraday cup configuration used for osmium isotope measurements by static multi-collection on a Thermo
 Neptune MC-ICP-MS and Thermo Triton Plus N-TIMS at Durham University. Only principal ions measured are listed,
 see ref. 37 (Luguet et al., 2008) for a full list of known polyatomic interferences.

Table 3 Osmium stable isotope compositions ($\delta^{190/188}Os$) of analytical tests on the MC-ICP-MS showing the robustness of our methodology with regards to (i) signal intensity and (ii) matrix effects and interference corrections. The percentages indicate the presence of each named element at x% of the Os abundance, e.g. Si was present at 5 % of the Os abundance. *Lith = Lithophile elements (Si, Al, Mg, Fe, Ca, Ti, Ni, and Cr); **PGE = Platinum Group Elements (Ru, Rh, Pd, Ir, and Pt).

Table 4 Average osmium stable and radiogenic isotope ratios of DROsS, SpecPure, ROMIL and OsCaR obtained in
 this study using: the double spike (DS) method; sample-standard bracketing by MC-ICP-MS, using non DS solutions
 and DROsS as the reference bracketing standard; and un-spiked runs by N-TIMS. For comparison the radiogenic
 isotope composition of DROsS as obtained by previous studies is included (Nowell et al., 2008³⁶; Luguet et al.,
 2008³⁷). Note that the better reproducibility (2 sd) for MC-ICP-MS relative to N-TIMS is, predominantly, related to
 difference in beam intensities at which analyses have been made (11-18 V vs. 1-18 V total Os beam, respectively).

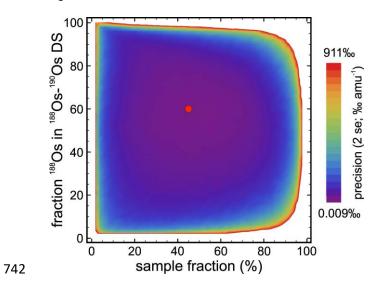
733 * - Four individual sample-standard bracketing sessions were run, on different days, and include 16 analyses all

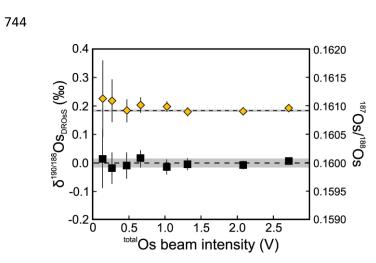
together. Presented are the average and 2 sd on the averages of the four analytical sessions.

735**Table 5** Osmium stable ($\delta^{190/188}Os$) and radiogenic ($^{187}Os/^{188}Os$ and $^{186}Os/^{188}Os$) isotope compositions, and Os736abundances of geological materials obtained by either MC-ICP-MS or N-TIMS. Samples are digested using carius737tubes (CT) or the high-pressure asher (HPA) system. W. mean = weighted mean; * = aliquots of sample material738that are combined after digestion and prior to chemical separation of Os.



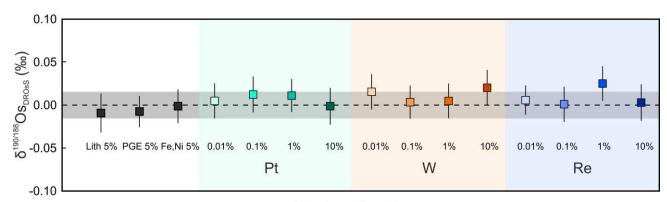
741 Figure 1







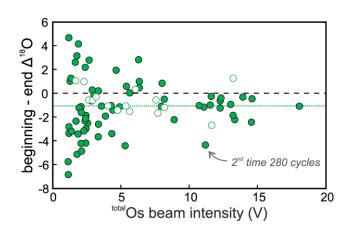


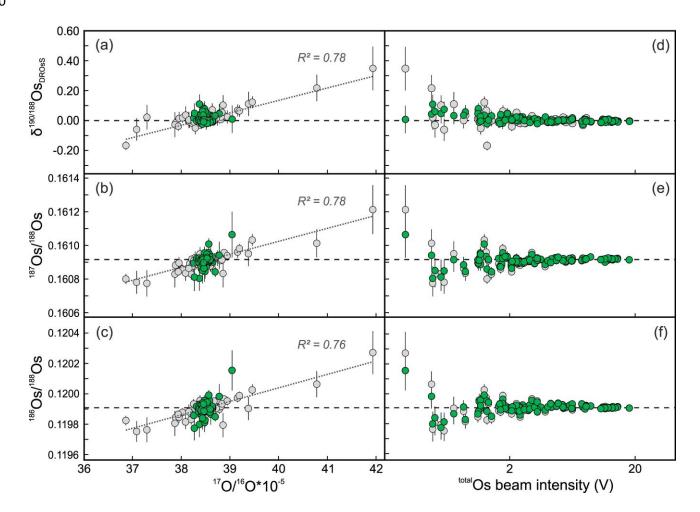


Number of analyses

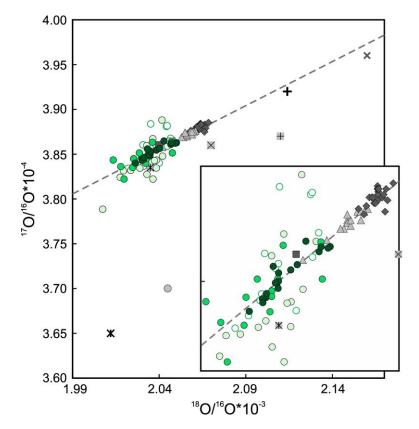












This study

In-run oxygen (intensity on mass 241)

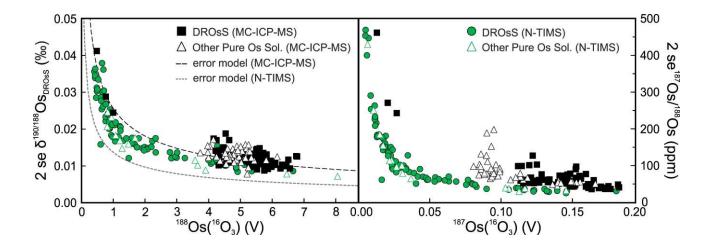
- >10 mV
- 5-10 mV
- <5 mV
- Pre and post oxygen

Previous studies

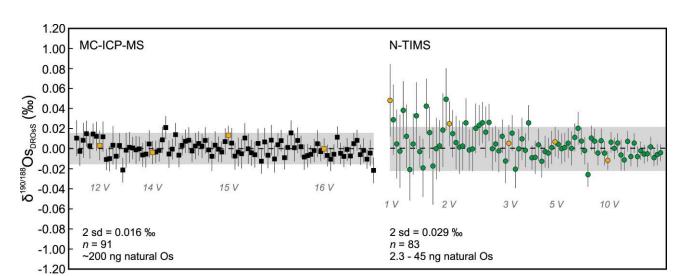
- Nagai and Yokoyama (2016)
- △ Worsham et al. (2016)
- X Luguet et al. (2008)
- Griselin et al. (2001)
- X Chavagnac (1998)
- + Thirlwall (1991)
- **X** Reisberg and Zindler (1986)
- × Nyquist in Wasserburg et al. (1981)
- Nier (1950)

---- TFL (slope 0.526; intercept 0.0099)

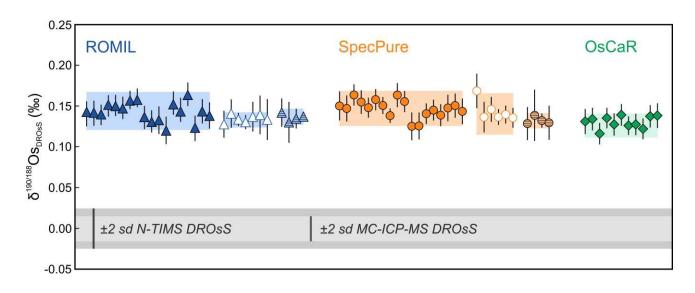
753 Figure 7



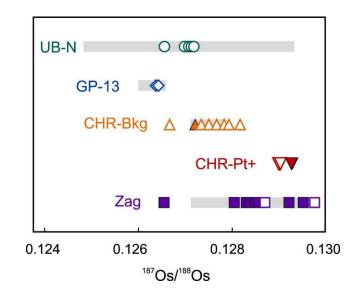




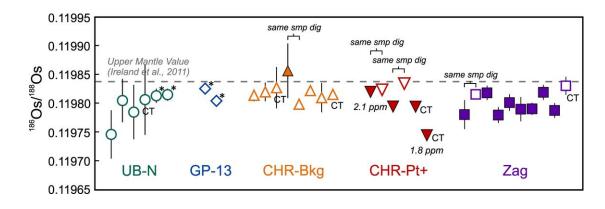




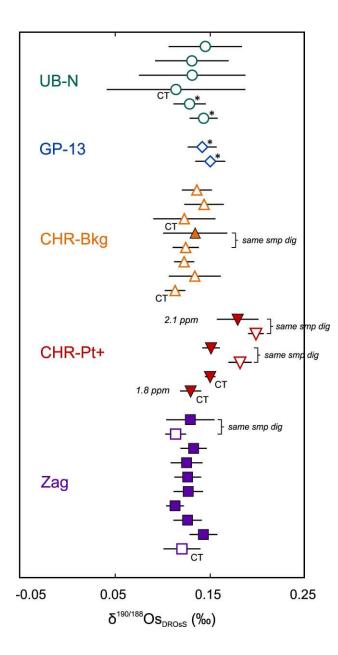
```
759 Figure 10
```



761 Figure 11

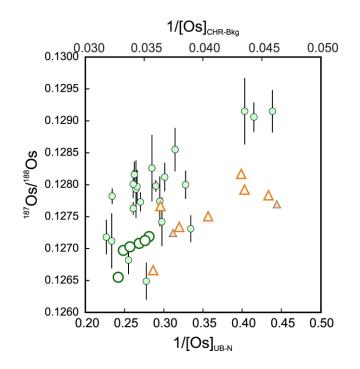






765 Figure 13





768 TABLES

769 Table 1

	¹⁸⁶ Os ¹⁸⁸ Os	¹⁸⁷ Os/ ¹⁸⁸ Os	¹⁸⁹ Os/ ¹⁸⁸ Os	¹⁹⁰ Os/ ¹⁸⁸ Os	¹⁹² Os/ ¹⁸⁸ Os
Reference std DROsS (±2 sd)	0.119909 ± 04	0.160916 ± 04	1.219780	1.983979 ± 30	3.083580 ± 1
¹⁸⁸ Os - ¹⁹⁰ Os spike (±2 se)	0.001372 ± 54	0.001239 ± 58	0.020944 ± 11	0.684512 ± 06	0.036517 ± 14

Cup	L4	L3	L2	L1	Ах	H1	H2	H3	H4
MC-ICP-MS									
Analyte isotopes	¹⁸³ W	¹⁸⁵ Re	¹⁸⁶ Os	¹⁸⁷ Os	¹⁸⁸ Os	¹⁸⁹ Os	¹⁹⁰ Os	¹⁹² Os	¹⁹⁴ Pt
N-TIMS									
Mass		232	234	235	236	237	238	240	
Seq1		¹⁸⁴ Os ¹⁶ O ₃ ⁻	¹⁸⁶ Os ¹⁶ O ₃ ⁻	¹⁸⁷ Os ¹⁶ O ₃ ⁻	¹⁸⁸ Os ¹⁶ O ₃ ⁻	¹⁸⁹ Os ¹⁶ O ₃ ⁻	¹⁹⁰ Os ¹⁶ O ₃ ⁻	¹⁹² Os ¹⁶ O ₃ ⁻	
Mass		236	238	239	240	241	242		
Seq2		¹⁸⁸ Os ¹⁶ O ₃ ⁻	¹⁹⁰ Os ¹⁶ O ₃ ⁻		¹⁹² Os ¹⁶ O ₃ ⁻	¹⁹² Os ¹⁶ O ₂ ¹⁷ O ⁻	¹⁹² Os ¹⁶ O ₂ ¹⁸ O ⁻		

(i) Signal intensity			
Concentration	δ ¹⁹⁰ Os	2 se	^{Total} Os (V)
0.10 ppm	0.007	0.012	2.71
0.08 ppm	-0.007	0.014	2.08
0.05 ppm	-0.004	0.021	1.31
0.04ppm	-0.014	0.026	1.02
0.03 ppm	0.017	0.033	0.66
0.02 ppm	-0.009	0.045	0.47
0.01 ppm	-0.018	0.055	0.26
0.005 ppm	0.014	0.102	0.14
(ii) Matrix effects and in	nterference cor	rections	
Doping element	δ ¹⁹⁰ Os	2 se	
Lith* 5%	-0.009	0.018	
PGE** 5%	-0.008	0.014	
Fe, Ni 5%	-0.001	0.017	
Pt 0.01%	0.005	0.017	
Pt 0.1%	0.012	0.018	
Pt 1%	0.011	0.016	
Pt 10%	-0.001	0.017	
W 0.01%	0.016	0.017	
W 0.1%	0.003	0.015	
W 1%	0.005	0.017	
W 10%	0.020	0.016	
Re 0.01%	0.006	0.012	
Re 0.1%	0.001	0.013	
Re 1%	0.025	0.014	
Re 10%	0.003	0.017	

	n	δ ¹⁹⁰ Os	2 sd	¹⁸⁷ Os/ ¹⁸⁸ Os	2 sd	¹⁸⁶ Os/ ¹⁸⁸ Os	2 sd
DROsS							
DS MC-ICP-MS	91	0.000	0.016	0.160916	0.000020	0.119909	0.000017
DS N-TIMS	83	0.001	0.029	0.160916	0.000043	0.119909	0.000043
un-spiked	14	-	-	0.160919	0.000031	-	-
previous study (MC-ICP-MS; Nowell et al., 2008 ³⁵)	21	-	-	0.160924	0.000003	0.119920	0.000002
previous study (N-TIMS; Luguet et al. ,2008 ³⁶)	8	-	-	0.160924	0.000004	0.119932	0.000006
ROMIL							
DS MC-ICP-MS	18	0.144	0.023	0.106827	0.000047	0.119806	0.000017
DS N-TIMS	7	0.133	0.009	0.106857	0.000042	0.119803	0.000023
sample-std bracketing MC-ICP-MS	4*	0.137	0.010	0.106861	0.000024	0.119806	0.000007
un-spiked	3	-	-	0.106878	0.000032	-	-
SpecPure							
DS MC-ICP-MS	18	0.147	0.022	0.149162	0.000026	0.119890	0.000019
DS N-TIMS	6	0.140	0.025	0.149161	0.000053	0.119889	0.000031
sample-std bracketing MC-ICP-MS	4*	0.132	0.009	0.149163	0.000012	0.119885	0.000016
un-spiked	3	-	-	0.149180	0.000088	-	-
OsCaR							
DS MC-ICP-MS	11	0.126	0.015	0.128298	0.000022	0.119816	0.000014

				smp weight (g)	^{total} Os (V)	¹⁸⁸ Os(O ₃ -)	¹⁸⁷ Os (O ₃ -)	δ ¹⁹⁰ Os	2 se	¹⁸⁷ Os/ ¹⁸⁸ Os	2 se	¹⁸⁶ Os/ ¹⁸⁸ Os	2 se	[Os] ng g ⁻¹	2 se
UB-N	1	HPA	N-TIMS	2.51	0.98	0.37	0.008	0.145	0.039	0.126555	0.000040	0.119746	0.000042	4.13	0.002
	2	HPA	N-TIMS	2.51	1.16	0.43	0.009	0.131	0.039	0.126983	0.000039	0.119804	0.000037	4.02	0.002
	3	HPA	N-TIMS	2.49	0.92	0.35	0.007	0.131	0.056	0.127061	0.000047	0.119784	0.000047	3.89	0.003
	4	СТ	N-TIMS	2.49	0.82	0.33	0.006	0.114	0.073	0.127073	0.000057	0.119806	0.000061	3.72	0.002
	5*	HPA	N-TIMS	6.03	3.99	1.53	0.030	0.128	0.017	0.127128	0.000011	0.119813	0.000012	3.62	0.001
	6*	HPA	N-TIMS	5.99	3.09	1.19	0.023	0.143	0.015	0.127181	0.000011	0.119815	0.000010	3.55	0.001
w. mean								0.136		0.127126		0.119813		3.64	
2 sd								0.022		0.000453		0.000053		0.46	
GP-13	1*	HPA	N-TIMS	4.02	4.95	1.87	0.039	0.142	0.015	0.126372	0.000010	0.119825	0.000010	3.87	0.002
	2*	HPA	N-TIMS	4.05	4.67	1.76	0.036	0.150	0.016	0.126426	0.000009	0.119804	0.000009	3.82	0.002
w. mean								0.146		0.126401		0.119813		3.84	
2 sd								0.012		0.000077		0.000030		0.07	
CHR-Bkg	1	HPA	N-TIMS	0.93	4.68	1.75	0.037	0.136	0.016	0.126667	0.000010	0.119813	0.000010	27.96	0.01
	2	HPA	N-TIMS	0.92	2.74	1.03	0.022	0.144	0.021	0.127671	0.000016	0.119819	0.000016	27.48	0.01
	3	СТ	N-TIMS	0.55	1.21	0.46	0.009	0.123	0.033	0.127340	0.000034	0.119827	0.000035	26.32	0.01
	4a	HPA	MC-ICP-MS	2.06	3.42	1.29	0.026	0.134	0.034	0.127210	0.000040	0.119856	0.000047	24.72	0.001
	4b	HPA	N-TIMS		6.02	2.30	0.046	0.124	0.014	0.127503	0.000008	0.119798	0.000009	24.72	0.001
	5	HPA	N-TIMS	0.92	9.64	3.80	0.070	0.123	0.011	0.128170	0.000006	0.119822	0.000006	23.11	0.01
	6	СТ	N-TIMS	0.58	1.81	0.71	0.013	0.134	0.027	0.127929	0.000025	0.119809	0.000025	22.97	0.01
	7	HPA	N-TIMS	0.92	9.50	3.79	0.067	0.113	0.011	0.127830	0.000006	0.119815	0.000007	21.95	0.01
w. mean								0.124		0.127803		0.119818		24.73	
2 sd								0.020		0.000945		0.000034		4.39	
CHR-Pt+	1a	HPA	MC-ICP-MS	0.23	15.54	5.62	0.129	0.179	0.022	0.128992	0.000011	0.119819	0.000009	2121	6
	1b	HPA	N-TIMS		15.35	5.59	0.129	0.198	0.008	0.129278	0.000004	0.119823	0.000005	2121	4
	2a	HPA	MC-ICP-MS	0.25	25.05	9.41	0.195	0.151	0.009	0.129012	0.000006	0.119794	0.000004	1805	6
	2b	HPA	N-TIMS		6.60	2.49	0.052	0.181	0.012	0.129260	0.000008	0.119833	0.000008	1805	6

	3	СТ	MC-ICP-MS	0.25	23.15	8.76	0.178	0.150	0.005	0.129029	0.000004	0.119796	0.000003	1785	4
	4	СТ	MC-ICP-MS	0.25	12.02	4.54	0.092	0.129	0.011	0.128993	0.000012	0.119744	0.000009	1747	6
w. mean								0.162		0.129033		0.119804		1879	
2 sd								0.051		0.000273		0.000064		349	
ZAG	1a	HPA	MC-ICP-MS	0.24	6.73	2.51	0.049	0.129	0.025	0.129518	0.000018	0.119780	0.000025	929	6
	1b	HPA	N-TIMS		3.28	1.26	0.025	0.114	0.011	0.129759	0.000007	0.119815	0.000007	929	4
	2	HPA	MC-ICP-MS	0.25	8.94	3.28	0.067	0.133	0.014	0.126549	0.000011	0.119818	0.000012	926	6
	3	HPA	MC-ICP-MS	0.25	9.16	3.38	0.068	0.125	0.017	0.128046	0.000014	0.119779	0.000013	905	6
	4	HPA	MC-ICP-MS	0.25	8.63	3.18	0.064	0.126	0.014	0.128557	0.000012	0.119801	0.000014	904	6
	5	HPA	MC-ICP-MS	0.50	5.78	2.15	0.042	0.127	0.016	0.129214	0.000028	0.119789	0.000021	892	3
	6	HPA	MC-ICP-MS	0.50	12.05	4.47	0.088	0.113	0.009	0.128650	0.000011	0.119790	0.000010	887	3
	7	HPA	MC-ICP-MS	0.25	8.73	3.24	0.064	0.126	0.015	0.128382	0.000013	0.119819	0.000013	884	6
	8	HPA	MC-ICP-MS	0.25	9.03	3.40	0.065	0.143	0.015	0.128303	0.000014	0.119787	0.000013	861	6
	9	СТ	N-TIMS	0.08	5.05	1.90	0.040	0.120	0.019	0.128699	0.000015	0.119830	0.000015	839	17
w. mean								0.123		0.128407		0.119804		896	
2 sd								0.018		0.001791		0.000036		59	
0															

779	REFER	ENCES
780	1.	S. B. Shirey and R. J. Walker, Annual Review of Earth and Planetary Sciences, 1998, 26, 423-
781		500.
782	2.	R. W. Carlson, <i>Lithos</i> , 2005, 82 , 249-272.
783	3.	R. J. Walker, M. F. Horan, J. W. Morgan, H. Becker, J. N. Grossman and A. E. Rubin, Geochimica
784		et Cosmochimica Acta, 2002, 66 , 4187-4201.
785	4.	T. Meisel, R. J. Walker and J. W. Morgan, <i>Nature</i> , 1996, 383 , 517-520.
786	5.	A. D. Brandon, R. J. Walker, J. W. Morgan, M. D. Norman and H. M. Prichard, <i>Science</i> , 1998,
787 788	6.	280 , 1570-1573. R. J. Walker, <i>Chemie der Erde-Geochemistry</i> , 2009, 69 , 101-125.
789	о. 7.	J. E. Snow and L. Reisberg, Earth and Planetary Science Letters, 1995, 133 , 411-421.
790	7. 8.	A. Meibom, N. H. Sleep, C. P. Chamberlain, R. G. Coleman, R. Frei, M. T. Hren and J. L. Wooden,
791	0.	Nature, 2002, 419 , 705-708.
792	9.	A. D. Brandon, R. A. Creaser, S. B. Shirey and R. W. Carlson, <i>Science</i> , 1996, 272 , 861-863.
793	10.	D. G. Pearson, S. W. Parman and G. M. Nowell, <i>Nature</i> , 2007, 449 , 202-205.
794	11.	C. W. Dale, A. Gannoun, K. W. Burton, T. W. Argles and I. J. Parkinson, Earth and Planetary
795		Science Letters, 2007, 253 , 211-225.
796	12.	H. J. Stein, K. Sundblad, R. J. Markey, J. W. Morgan and G. Motuza, Mineralium Deposita, 1998,
797		33 , 329-345.
798	13.	K. W. Burton, Journal of Geochemical Exploration, 2006, 88 , 262-265.
799	14.	W. J. Pegram, S. Krishnaswami, G. E. Ravizza and K. K. Turekian, Earth and Planetary Science
800		<i>Letters</i> , 1992, 113 , 569-576.
801 802	15.	B. Peucker-Ehrenbrink and G. Ravizza, <i>Terra Nova</i> , 2000, 12 , 205-219.
802 803	16.	S. G. Nielsen, M. Rehkämper and A. N. Halliday, <i>Geochimica et Cosmochimica Acta</i> , 2006, 70 , 2643-2657.
803 804	17.	G. A. Brennecka, L. E. Borg, I. D. Hutcheon, M. A. Sharp and A. D. Anbar, <i>Earth and Planetary</i>
805	17.	Science Letters, 2010, 291 , 228-233.
806	18.	J. B. Creech, J. A. Baker, M. R. Handler, M. Schiller and M. Bizzarro, <i>Journal of Analytical Atomic</i>
807		Spectrometry, 2013, 28 , 853-865.
808	19.	J. B. Creech, J. A. Baker, M. R. Handler and M. Bizzarro, <i>Chemical Geology</i> , 2014, 363 , 293-300.
809	20.	C. H. Stirling, M. B. Andersen, EK. Potter and A. N. Halliday, Earth and Planetary Science
810		Letters, 2007, 264 , 208-225.
811	21.	M. Rehkämper, M. Frank, J. R. Hein, D. Porcelli, A. Halliday, J. Ingri and V. Liebetrau, <i>Earth and</i>
812		Planetary Science Letters, 2002, 197 , 65-81.
813	22.	S. G. Nielsen, M. Rehkamper, M. D. Norman, A. N. Halliday and D. Harrison, <i>Nature</i> , 2006, 439 ,
814	22	314-317.
815 816	23.	J. J. Shen, D. A. Papanastassiou and G. J. Wasserburg, <i>Geochimica et Cosmochimica Acta</i> , 1996,
810	24.	60 , 2887-2900. J. L. Birck, M. R. Barman and F. Capmas <i>, Geostandards Newsletter</i> , 1997, 21 , 19-27.
818	2 4 . 25.	M. H. Dodson, Journal of Scientific Instruments, 1963, 40 , 289-295.
819	26.	S. J. G. Galer, <i>Chemical Geology</i> , 1999, 157 , 255-274.
820	27.	MA. Millet, J. A. Baker and C. E. Payne, <i>Chemical Geology</i> , 2012, 304 , 18-25.
821	28.	A. D. Brandon, M. Humayun, I. S. Puchtel, I. Leya and M. Zolensky, Science, 2005, 309, 1233-
822		1236.
823	29.	T. Yokoyama, V. K. Rai, C. M. D. Alexander, R. S. Lewis, R. W. Carlson, S. B. Shirey, M. H.
824		Thiemens and R. J. Walker, <i>Earth and Planetary Science Letters</i> , 2007, 259 , 567-580.
825	30.	T. Yokoyama, C. M. D. Alexander and R. J. Walker, Earth and Planetary Science Letters, 2010,
826	~ ~	291 , 48-59.
827 020	31.	N. Wittig, M. Humayun, A. D. Brandon, S. Huang and I. Leya, <i>Earth and Planetary Science</i>
828 829	22	Letters, 2013, 361 , 152-161.
829 830	32.	C. Siebert, T. F. Nägler and J. D. Kramers, <i>Geochemistry, Geophysics, Geosystems</i> , 2001, 2 , 1032.
830	33.	MA. Millet and N. Dauphas, <i>Journal of Analytical Atomic Spectrometry</i> , 2014, 29 , 1444-1458.
832	33. 34.	MA. Millet, N. Dauphas, N. D. Greber, K. W. Burton, C. W. Dale, B. Debret, C. G. Macpherson,
833		G. M. Nowell and H. M. Williams, <i>Earth and Planetary Science Letters</i> , 2016, 449 , 197-205.
834	35.	J. F. Rudge, B. C. Reynolds and B. Bourdon, <i>Chemical Geology</i> , 2009, 265 , 420-431.

835	36.	G. M. Nowell, A. Luguet, D. G. Pearson and M. A. Horstwood, Chemical Geology, 2008, 248,
836		363-393.
837	37.	A. Luguet, G. M. Nowell and D. G. Pearson, Chemical Geology, 2008, 248, 342-362.
838	38.	J. Liu and D. G. Pearson, Chemical Geology, 2014, 363, 301-311.
839	39.	R. Chatterjee and J. C. Lassiter, Chemical Geology, 2015, 396 , 112-123.
840	40.	T. Meisel, L. Reisberg, J. Moser, J. Carignan, F. Melcher and G. Brügmann, Chemical geology,
841		2003, 201 , 161-179.
842	41.	D. G. Pearson, G. J. Irvine, D. A. Ionov, F. R. Boyd and G. E. Dreibus, Chemical Geology, 2004,
843		208 , 29-59.
844	42.	P. J. Potts, C. J. B. Gowing and K. Govindaraju, Geostandards Newsletter, 1992, 16, 81-108.
845	43.	V. Paliulionyte, T. Meisel, P. Ramminger and P. Kettisch, Geostandards and Geoanalytical
846		Research, 2006, 30 , 87-96.
847	44.	T. Meisel, J. Moser, N. Fellner, W. Wegscheider and R. Schoenberg, Analyst, 2001, 126, 322-
848		328.
849	45.	S. B. Shirey and R. J. Walker, Analytical Chemistry, 1995, 67, 2136-2141.
850	46.	A. S. Cohen and F. G. Waters, Analytica Chimica Acta, 1996, 332 , 269-275.
851	47.	M. Roy-Barman, Thesis, Université de Paris VII, 1993.
852	48.	C. W. Dale, K. W. Burton, D. G. Pearson, A. Gannoun, O. Alard, T. W. Argles and I. J. Parkinson,
853		Geochimica et Cosmochimica Acta, 2009, 73 , 1394-1416.
854	49.	R. A. Creaser, D. A. Papanastassiou and G. J. Wasserburg, Geochimica et Cosmochimica Acta,
855		1991, 55 , 397-401.
856	50.	F. Albarède and B. Beard, Reviews in Mineralogy and Geochemistry, 2004, 55, 113-152.
857	51.	Y. Nagai and T. Yokoyama, Journal of Analytical Atomic Spectrometry, 2016, 31 , 948-960.
858	52.	E. A. Worsham, R. J. Walker and K. R. Bermingham, International Journal of Mass Spectrometry,
859		2016, 407 , 51-61.
860	53.	Y. Liu, M. Huang, A. Masuda and M. Inoue, International Journal of Mass Spectrometry and Ion
861		Processes, 1998, 173 , 163-175.
862	54.	ZY. Chu, CF. Li, Z. Chen, JJ. Xu, YK. Di and JH. Guo, Analytical chemistry, 2015.
863	55.	K. R. Bermingham, R. J. Walker and E. A. Worsham, International Journal of Mass Spectrometry,
864		2016, 403 , 15-26.
865	56.	A. Trinquier, M. Touboul and R. J. Walker, Analytical chemistry, 2016, 88, 1542-1546.
866	57.	R. N. Clayton, L. Grossman and T. K. Mayeda, Science, 1973, 182, 485-488.
867	58.	E. Barkan and B. Luz, Rapid Communications in Mass Spectrometry, 2005, 19, 3737-3742.
868	59.	D. Rumble, M. F. Miller, I. A. Franchi and R. C. Greenwood, <i>Geochimica et Cosmochimica Acta</i> ,
869		2007, 71 , 3592-3600.
870	60.	I. Ahn, J. I. Lee, M. Kusakabe and BG. Choi, <i>Geosciences Journal</i> , 2012, 16 , 7-16.
871	61.	R. J. Walker, A. D. Brandon, J. M. Bird, P. M. Piccoli, W. F. McDonough and R. D. Ash, Earth and
872		Planetary Science Letters, 2005, 230 , 211-226.
873	62.	D. G. Pearson and S. J. Woodland, Chemical Geology, 2000, 165, 87-107.
874	63.	L. P. Bédard and SJ. Barnes, <i>Geostandards and Geoanalytical Research</i> , 2004, 28 , 311-316.
875	64.	T. J. Ireland, R. J. Walker and A. D. Brandon, Geochimica et Cosmochimica Acta, 2011, 75, 4456-
876		4475.
877	65.	L. Reisberg and T. Meisel, Geostandards Newsletter, 2002, 26, 249-267.
878	66.	A. Ishikawa, R. Senda, K. Suzuki, C. W. Dale and T. Meisel, <i>Chemical Geology</i> , 2014, 384 , 27-46.
879	67.	R. A. Lord, H. M. Prichard and C. R. Neary, Transactions - Institution of Mining & Metallurgy,
880		Section B, 1994.
881	68.	M. Griselin, J. Van Belle, C. Pomies, P. Vroon, M. Van Soest and G. Davies, <i>Chemical Geology</i> ,
882		2001, 172 , 347-359.
883	69.	V. Chavagnac, Université Rennes 1, 1998.
884	70.	M. Thirlwall, <i>Chemical Geology</i> , 1991, 94 , 13-22.
885	71.	L. Reisberg and A. Zindler, <i>Earth and Planetary Science Letters</i> , 1986, 81 , 29-45.
886	72.	G. Wasserburg, S. Jacobsen, D. DePaolo, M. McCulloch and T. Wen, <i>Geochimica et</i>
887		Cosmochimica Acta, 1981, 45 , 2311-2323.
888	73.	A. O. Nier, <i>Physical Review</i> , 1950, 77 , 789.
889	74.	J. Harvey, C. W. Dale, A. Gannoun and K. W. Burton, <i>Geochimica et Cosmochimica Acta</i> , 2011,
890		75 , 5574-5596.
891	75.	I. S. Puchtel, M. Humayun and R. J. Walker, <i>Precambrian Research</i> , 2007, 158 , 119-137.

76. C. Marchesi, C. W. Dale, C. J. Garrido, D. G. Pearson, D. Bosch, J.-L. Bodinier, F. Gervilla and K.
Hidas, *Earth and Planetary Science Letters*, 2014, **400**, 33-44.

897 **Electronic Supplementary Information**

898 S.1. Oxygen corrections for N-TIMS analyses

899 An important step in the N-TIMS data reduction, is the correction for oxygen interferences. In this study, 900 we have applied a two-step correction, where in step one a starting oxygen composition is used in order 901 to determine the "true" oxygen isotope composition of the particular measurement. For the first step 902 we have used the compositions as determined by Luguet et al. (2008):

903
$${}^{17}\text{O}/{}^{16}\text{O}_{\text{Luguet}} = R_1 = 0.00038582$$
 (1)

904
$${}^{18}O/{}^{16}O_{Luguet} = R_2 = 0.00203486$$
 (2)

905 The intensities obtained on a certain mass (I_x in V) were corrected for the effects of variable isobaric 906 oxygen isotope interferences of the heavier oxygen isotopes (¹⁷O and ¹⁸O) following:

907	$O_1 = 3^* R_1$	(3)
908	$O_2 = 3^* R_1^2 + 3^* R_2$	(4)
909	$O_3 = R_1^3 + 6^* R_1^* R_2$	(5)
910	$O_4 = 3^* R_1^{2*} R_2 + 3^* R_2^2$	(6)
911	$O_5 = 3^* R_1^* R_2^2$	(7)
912	$O_6 = R_2^3$	(8)

913 Where O_i represent the difference in mass (i in amu) between the analyte mass of interest and 914 the lightest mass that could generate a potential oxide interference. For example, the ions collected on 915 mass 234 (1234) are predominantly representing ¹⁸⁶Os¹⁶Os⁻ (1¹⁸⁶) ions but also reflect oxygen complexes originating from ¹⁸⁴Os (¹⁸⁴Os¹⁶O¹⁷O₂⁻ and ¹⁸⁴Os¹⁶O₂¹⁸O⁻). In this case, the mass difference is 2 amu and 916 thus $O_i = O_2$. Using equations 3-8, the intensities of ${}^{j}Os^{16}O_3^{-1}$ ions (l^{j}) can be determined as follows: 917

918	$I^{184} = I_{232}$	(9)
919	$I^{186} = I_{234} - (I_{232} * O_2)$	(10)
920	$I^{187} = I_{235} - (I_{232}^* O_3) - (I_{234}^* O_1)$	(11)
921	$I^{188} = I_{236} - (I_{232}^*O_4) - (I_{234}^*O_2) - (I_{235}^*O_1)$	(12)
922	$I^{189} = I_{237} - (I_{232}^*O_5) - (I_{234}^*O_3) - (I_{235}^*O_2) - (I_{236}^*O_1)$	(13)
923	$I^{190} = I_{238} - (I_{232}^*O_6) - (I_{234}^*O_4) - (I_{235}^*O_3) - (I_{236}^*O_2) - (I_{237}^*O_1)$	(14)
924	$I^{192} = I_{240} - (I_{234}^*O_6) - (I_{235}^*O_5) - (I_{236}^*O_4) - (I_{237}^*O_3) - (I_{238}^*O_2)$	(15)

By obtaining the intensities on the various ${}^{j}Os^{16}O_{3}^{-}$ ions, masses 241 (${}^{192}Os^{16}O_{2}{}^{17}O^{-} = I^{192}17$) and 242 925 $(^{192}Os^{16}O_2^{18}O^- = I^{192}18)$ can be stripped from minor tri-oxide interferences as well, following: 926

927
$$I^{192}17 = I_{241} - (I_{235}^*O_6) - (I_{236}^*O_5) - (I_{237}^*O_4) - (I_{238}^*O_3)$$
(16)

928
$$I^{192}18 = I_{242} - (I_{236}^*O_6) - (I_{237}^*O_5) - (I_{238}^*O_4)$$
 (17)

929 Subsequently, these intensities were used to determine the "true" oxygen isotope compositions 930 $(^{17}O/^{16}O \text{ and } ^{18}O/^{16}O)$ of the individual cycle:

931
$${}^{17}\text{O}/{}^{16}\text{O}_{\text{true}} = R'_1 = (l^{192}17/l^{192})/3$$
 (18)

932
$${}^{18}O/{}^{16}O_{true} = R'_2 = (I^{192}18/I^{192})/3$$
 (19)

933 These oxygen compositions were then used to perform the "stripping" again (equations 3-15), now with 934 the "true" oxygen composition. This means that R_1 and R_2 in equations 3-8 are replaced by R'_1 and R'_2 . 935 In this study, we observed that the ¹⁷O/¹⁶O composition imposes a significant inaccuracy on the 936 stable Os isotope composition when total Os signal intensities dropped below 1 V. Therefore, instead 937 of accomplishing equations (16) and (18), the ¹⁷O/¹⁶O composition was based on the "true or measured" 938 $^{18}\text{O}/^{16}\text{O}$ composition as determined in equation (19). In order to calculate the $^{17}\text{O}/^{16}\text{O}$ from the 939 measured ¹⁸O/¹⁶O the relationship between the stable oxygen isotopes needs to be considered. In this 940 study, we assume that the oxygen isotopic compositions vary as a result of equilibrium mass-dependent 941 isotopic fractionation. The relationship among the three stable oxygen isotopes is exponential and can 942 be written as:

943
$$\delta^{17} O + 1 = \left(\delta^{18} O + 1\right)^{\lambda}$$
(20)

944 , where λ defines the slope of the line. By taking a logarithm, equation (20) can be linearized (Miller, 945 2002¹):

946
$$10^3 * \ln(\delta^{17}O/10^3 + 1) = \lambda * \left(10^3 * \ln(\delta^{18}O/10^3 + 1)\right)$$
(21)

947 To calculate the ¹⁷O/¹⁶O composition, equation (21) can be re-written:

948
$$\delta^{17}0 = e^{\lambda * \left(ln \left(\left(\delta^{18} O / 10^3 \right) + 1 \right) \right)} * 10^3$$
(22)

949 with,

950
$${}^{17}\text{O}/{}^{16}\text{O}_{true} = \left(\delta^{17}\text{O}/10^3 + 1\right) * {}^{17}\text{O}/{}^{16}\text{O}_{VSMOW}$$
 (23)

951 , where λ is 0.526, as we assume that the fractionation occurs along the "Terrestrial Fractionation Line" 952 (TFL²) for which the average literature values is 0.526 $^{e.g.3-5}$. The δ^{17} O and δ^{18} O use the VSMOW (Vienna Standard Mean Ocean Water) as reference values. Generally, the ¹⁸O/¹⁶O ratio of VSMOW is accepted 953 to be 0.0020052^6 whereas there is more debate about the ${}^{17}O/{}^{16}O$ ratio, with values ranging between 954 955 0.000380⁷ and 0.000384⁸. Using data of this study we obtain an intercept with zero at a value of 956 0.000382 which is within the range previously observed and, therefore, used in this study (Fig. S.1). To 957 determine this intercept we used analyses that were obtained at a beam intensity of >2 mV on mass 241. If these analyses are plotted in $10^3 \ln(1+\delta^{17}O/10^3)$ vs. $10^3 \ln(1+\delta^{18}O/10^3)$ space, following the 958 959 approach of Miller $(2002)^1$, a regression coefficient (λ) of 0.488±0.072 (95 c.i.) is obtained (Fig. S.1). 960 Isoplot 4 was used to calculate the slope as well as the error on the slope. When the selected analyses 961 are considered in ${}^{17}\text{O}/{}^{16}\text{O}$ vs. ${}^{18}\text{O}/{}^{16}\text{O}$ space an approximate linear relationship with a slope of 0.092 ± 962 0.014 (95 c.i.) and an intercept of 0.000197 ± 0.000027 (95 c.i.) is obtained. This is within error of the 963 values obtained when assuming a slope of 0.526 over the ¹⁸O/¹⁶O interval of 0.002007 to 0.002055 (the 964 range in ${}^{18}O/{}^{16}O$ observed in this study); 0.0994 x + 0.000183, where x is ${}^{18}O/{}^{16}O$.

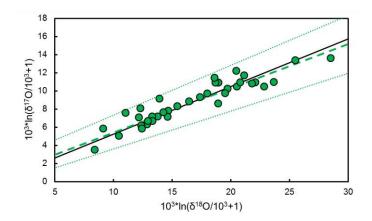


Figure S.1 Measured ¹⁷O/¹⁶O and ¹⁸O/¹⁶O compositions of analyses with a beam intensity of >2 mV on mass 241 plotted as delta values in logarithmic form. The regression line (dashed line) is within error (dotted lines) of the terrestrial fractionation line with a slope of 0.526 (solid line).

975

976 **REFERENCES**

- 977 1. M. F. Miller, *Geochimica et Cosmochimica Acta*, 2002, **66**, 1881-1889.
- 978 2. R. N. Clayton, L. Grossman and T. K. Mayeda, *Science*, 1973, **182**, 485-488.
- 979 3. E. Barkan and B. Luz, *Rapid Communications in Mass Spectrometry*, 2005, **19**, 3737-3742.
- D. Rumble, M. F. Miller, I. A. Franchi and R. C. Greenwood, *Geochimica et Cosmochimica Acta*, 2007, **71**, 3592-3600.
- 982 5. I. Ahn, J. I. Lee, M. Kusakabe and B.-G. Choi, *Geosciences Journal*, 2012, **16**, 7-16.
- 983 6. P. Baertschi, *Earth and Planetary Science Letters*, 1976, **31**, 341-344.
- 984 7. W. J. Li, B. L. Ni, D. Q. Jin and Q. G. Zhang, *Kexue Tongbao*, 1988, **33**, 1610-1613.
- 985 8. J. C. Lorin, *Chemical Geology: Isotope Geoscience section*, 1992, **101**, 193-195.

986